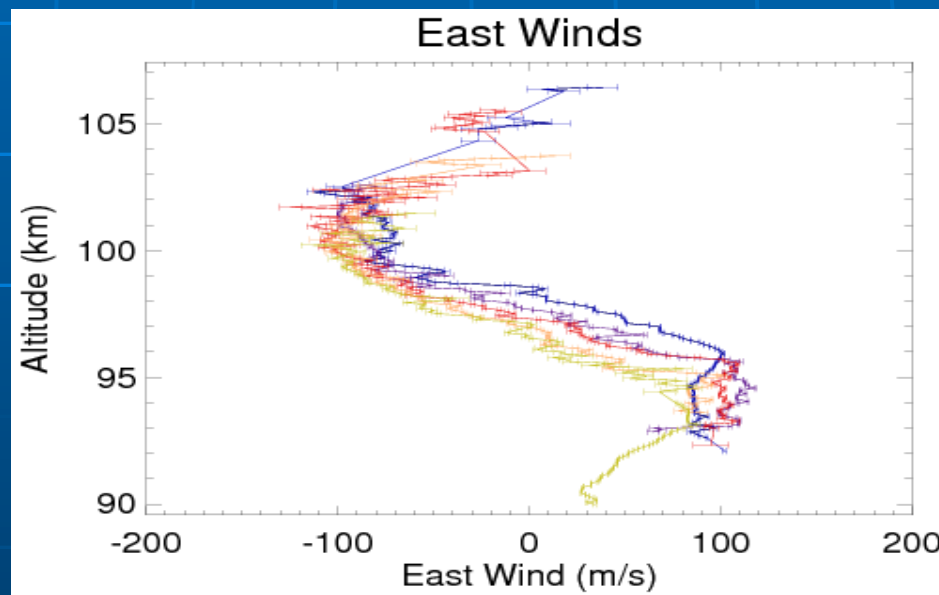


Measuring MLT Winds by Tracking Meteor Trails with Incoherent Scatter Radar

A New Method Observes Rapidly Changing Winds Exceeding 500km/Hr

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Jorge Chau², and Sigrid Close³

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Research supported
by the NSF
Aeronomy, DOE/NSF
& CEDAR Grants

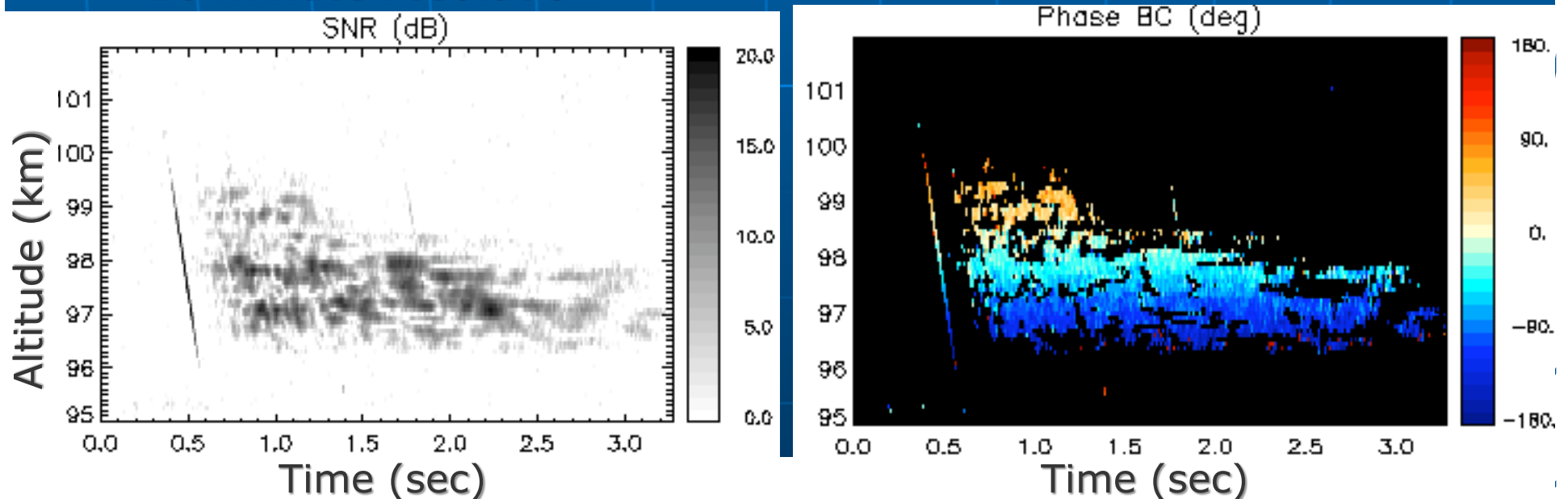
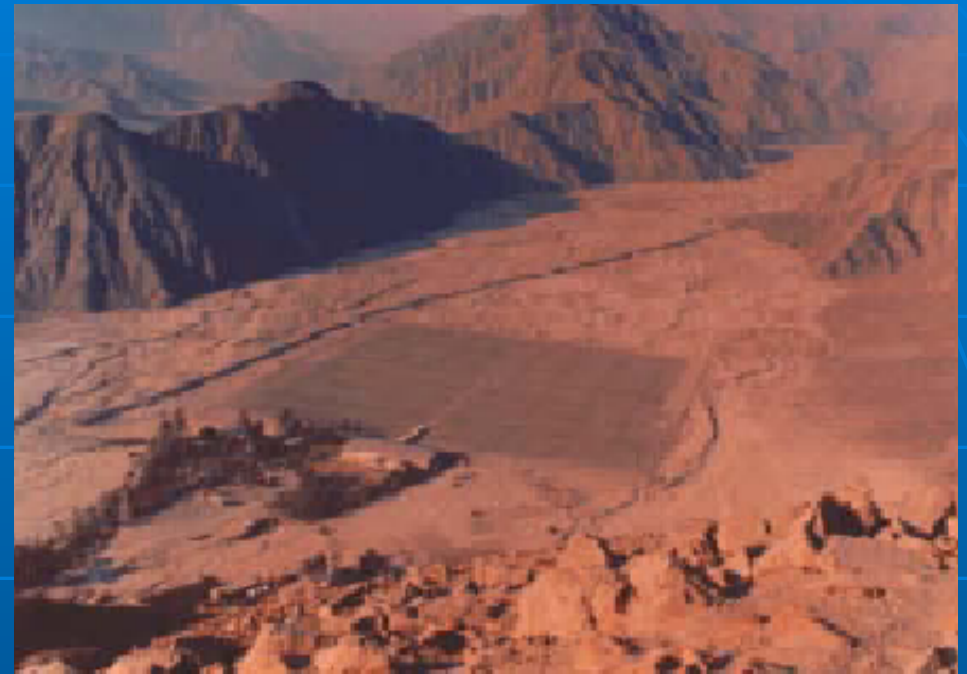
BOSTON
UNIVERSITY

CEDAR Workshop 2009



Overview

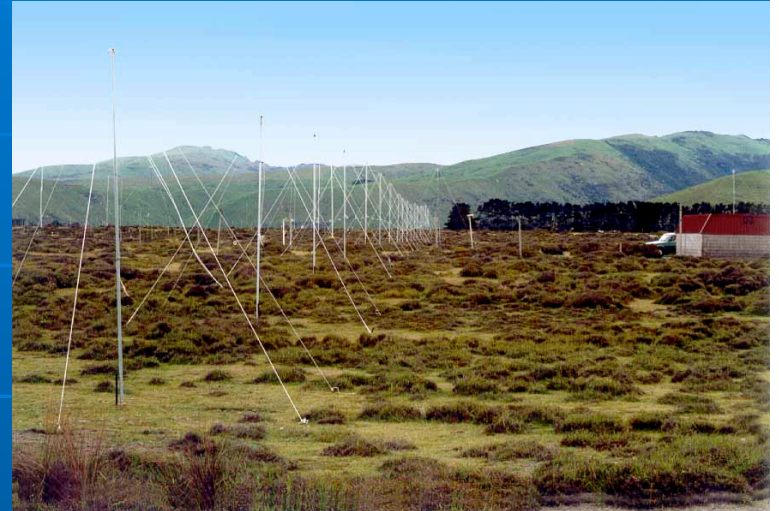
- Background:
 - Winds in the MLT
 - Observations of Specular Meteor trails by Large (HPLA) Radars
- Method: Use Phase interferometry from Trails
- Result: Accurate Wind measurements
 - km resolution
 - 10 minute resolution



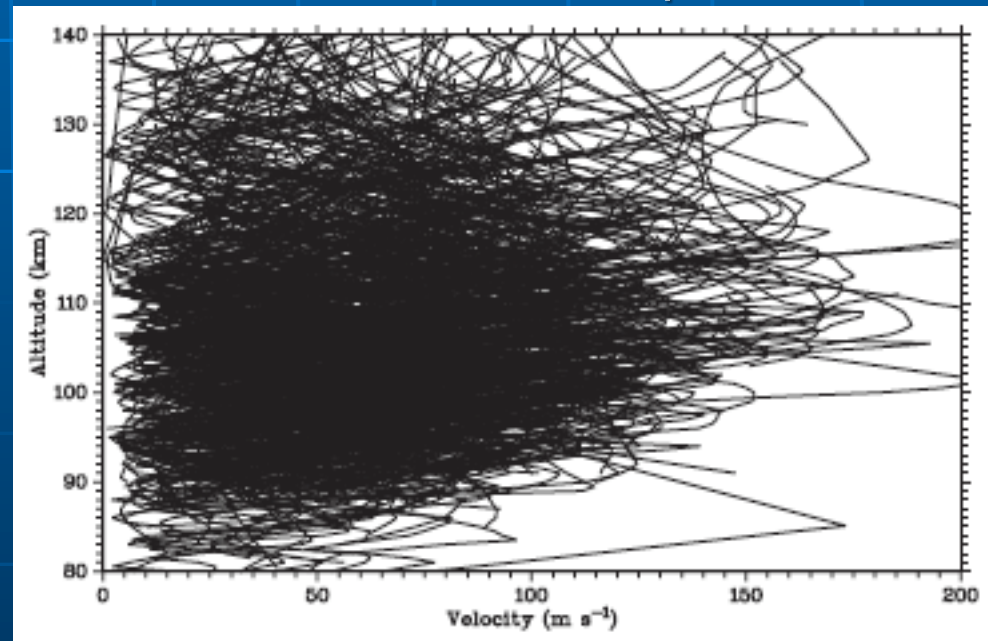
Winds in the Lower Thermosphere

- Specular Radar measurements
- LIDAR measurements
- 400 Rocket launches to measure winds

Specular Meteor Radar



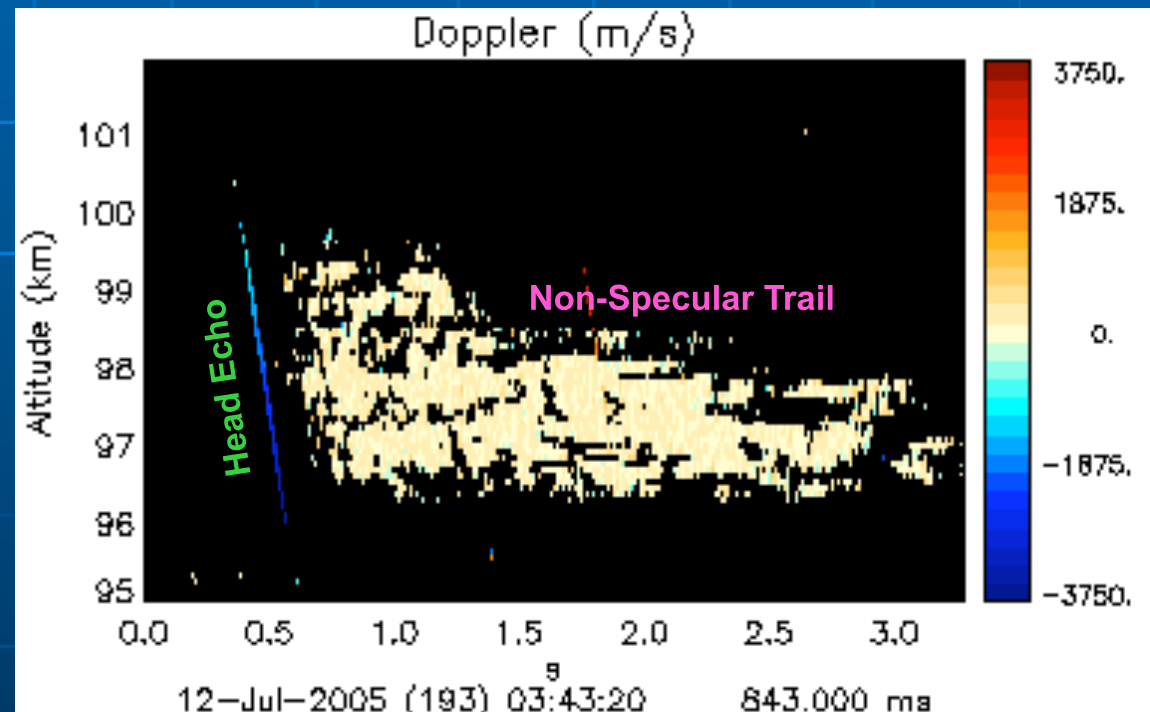
Rocket Wind Speeds



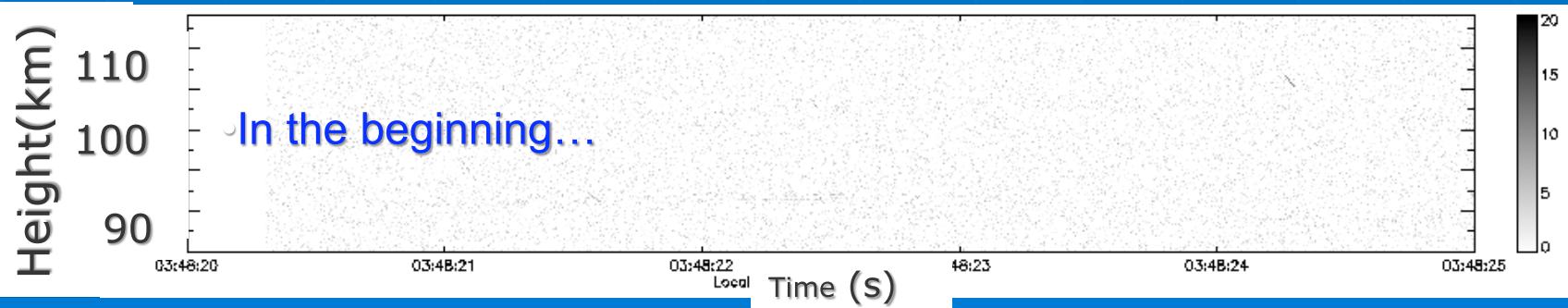
Background: HPLA Meteor Radar Science

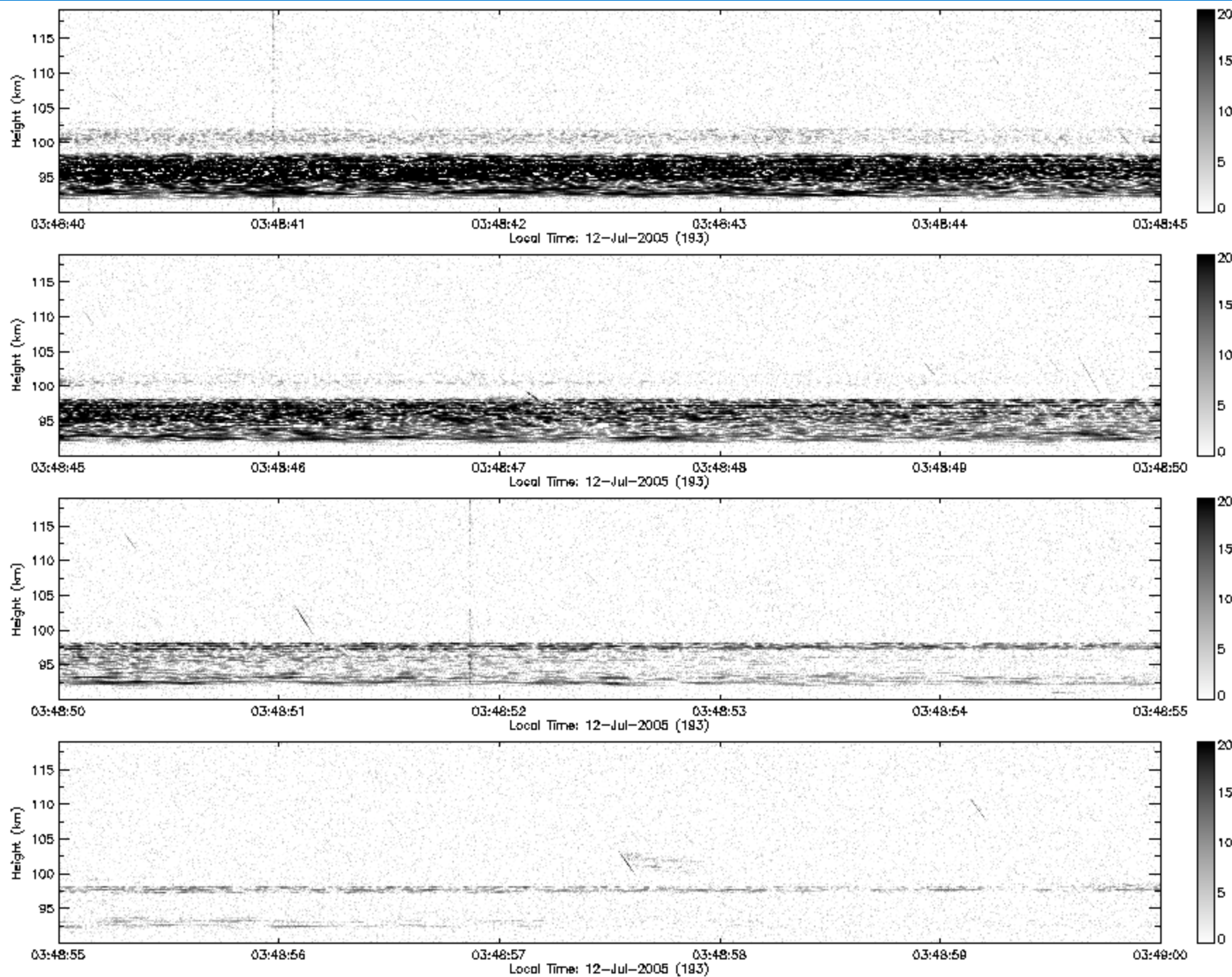


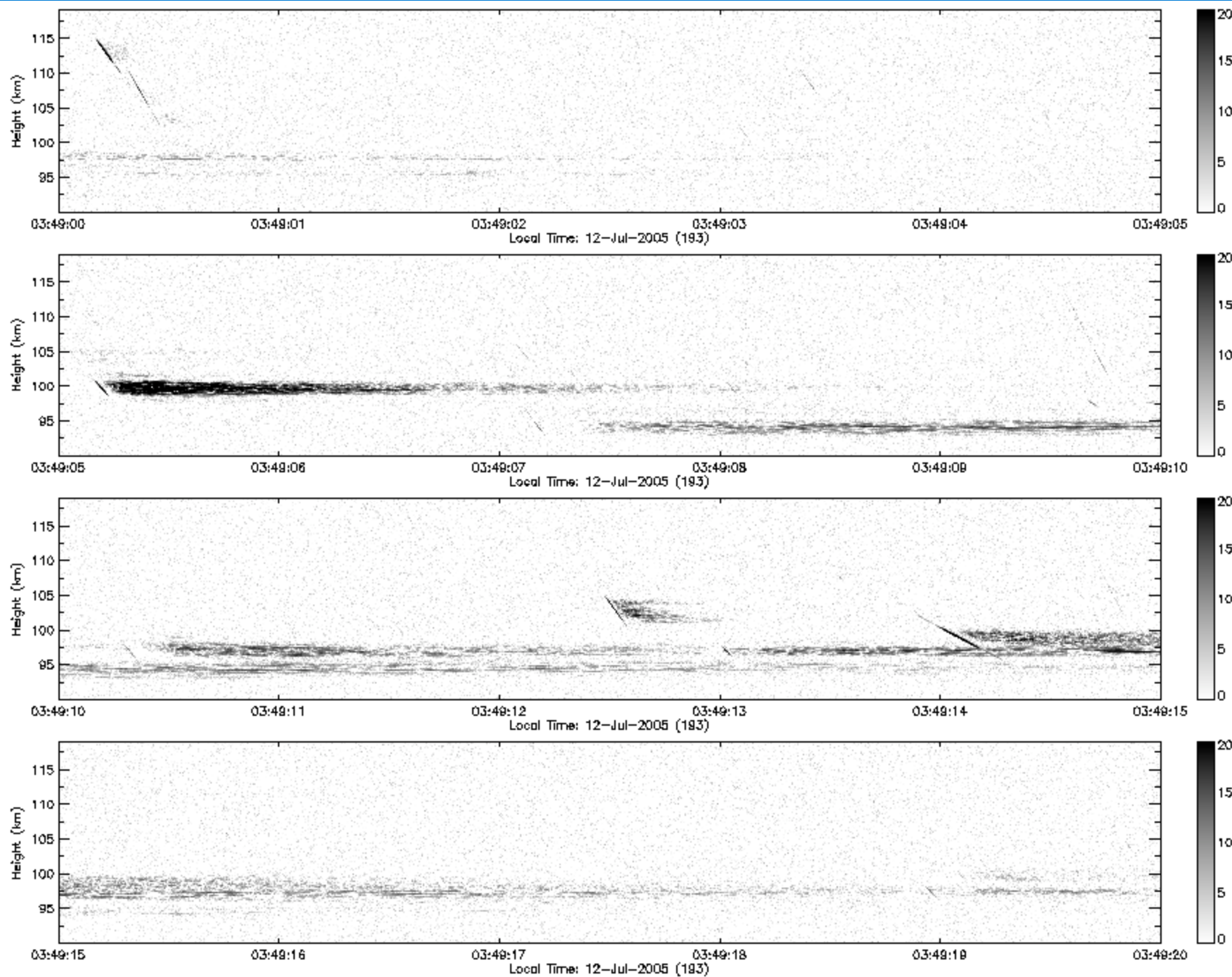
- Meteor enters Atmosphere
- Slam into air molecules
- Heat, ablate & ionize
- Create plasma trail
- Reflections from leading edge of plasma form head echoes
- Trail cools and expands
- Diffusion begins
- Waves and Turbulence develop:
 - In a limited range of altitudes
 - Reflections from plasma waves creates **non-specular trails**



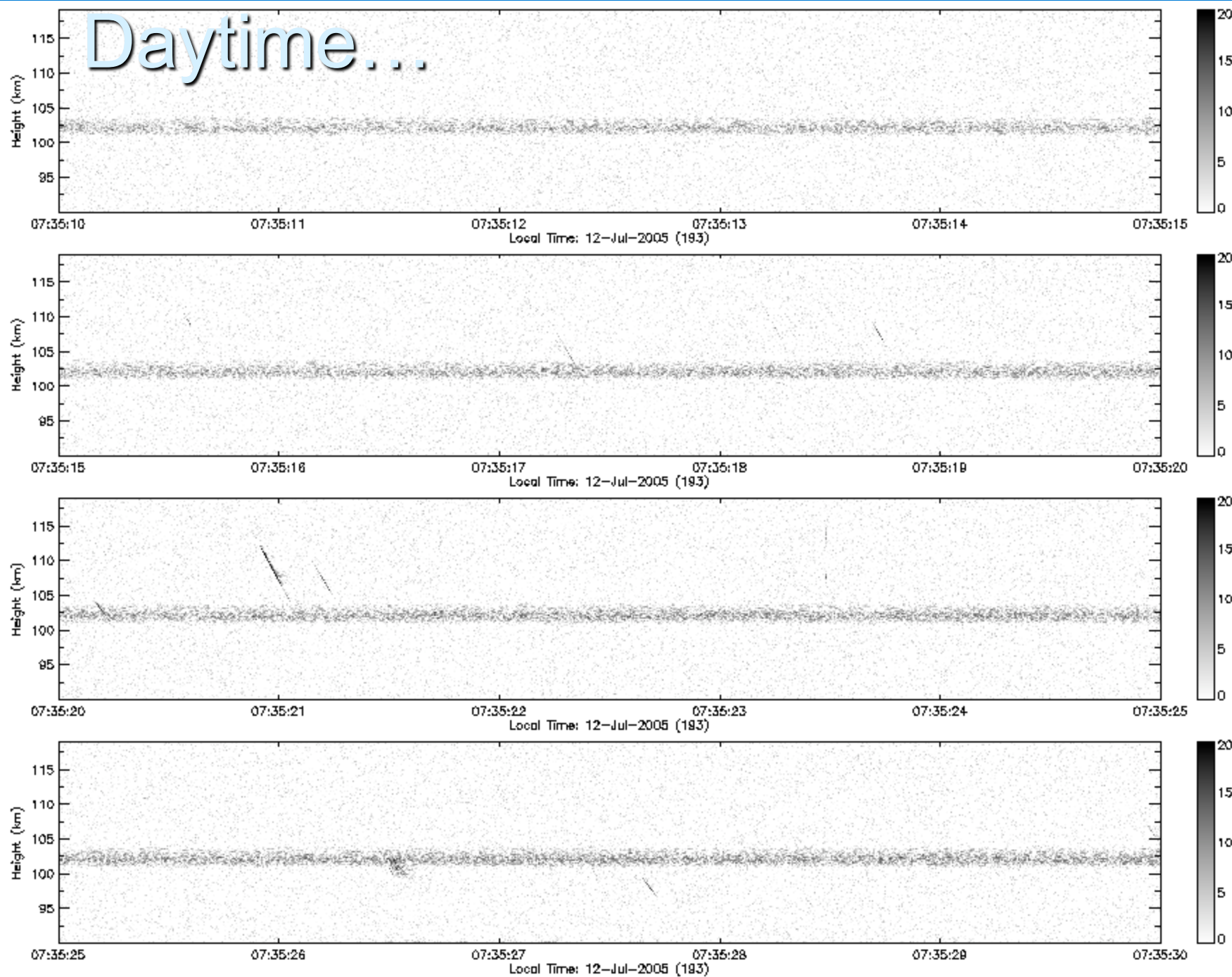
Data from JRO: My first 20 seconds

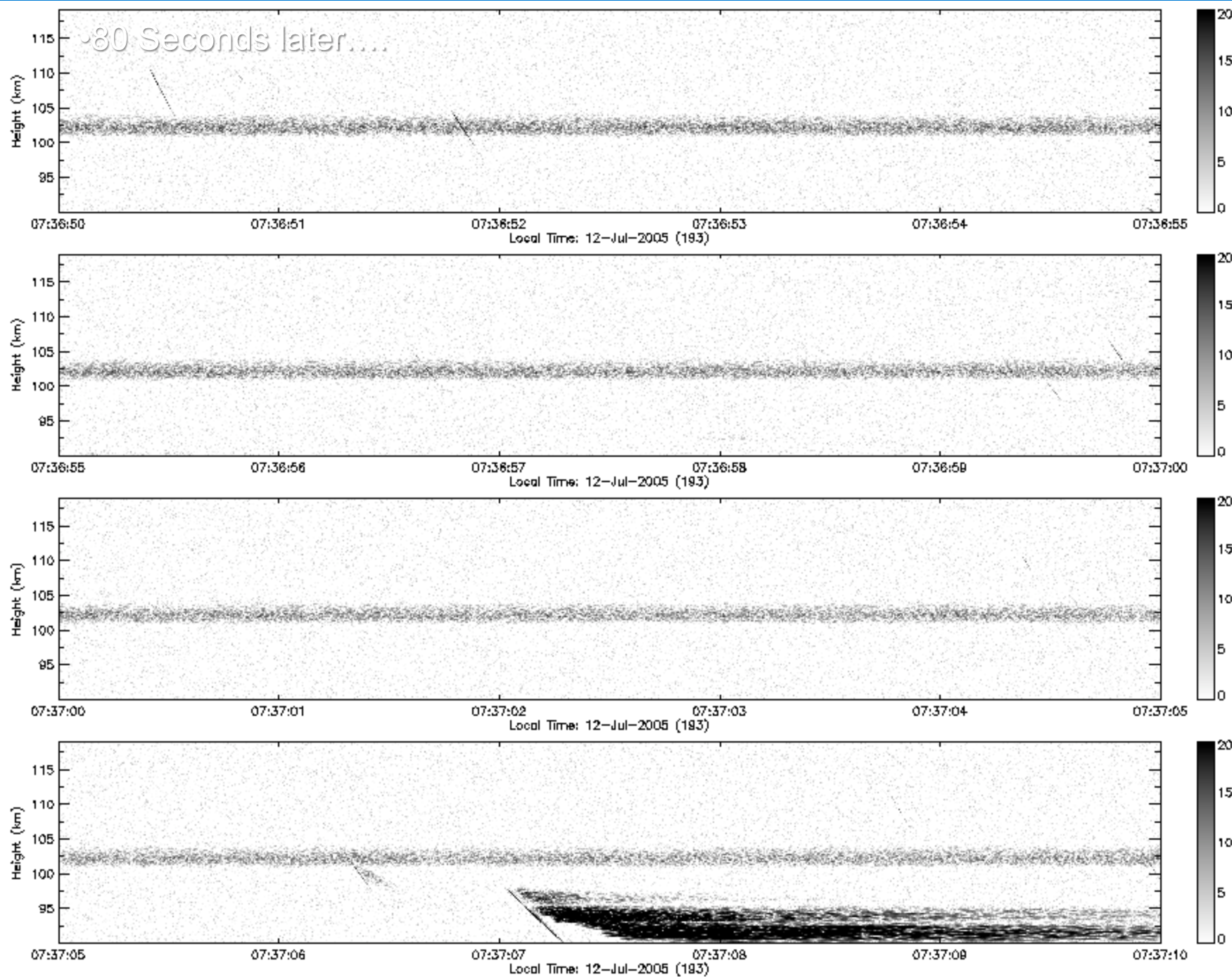




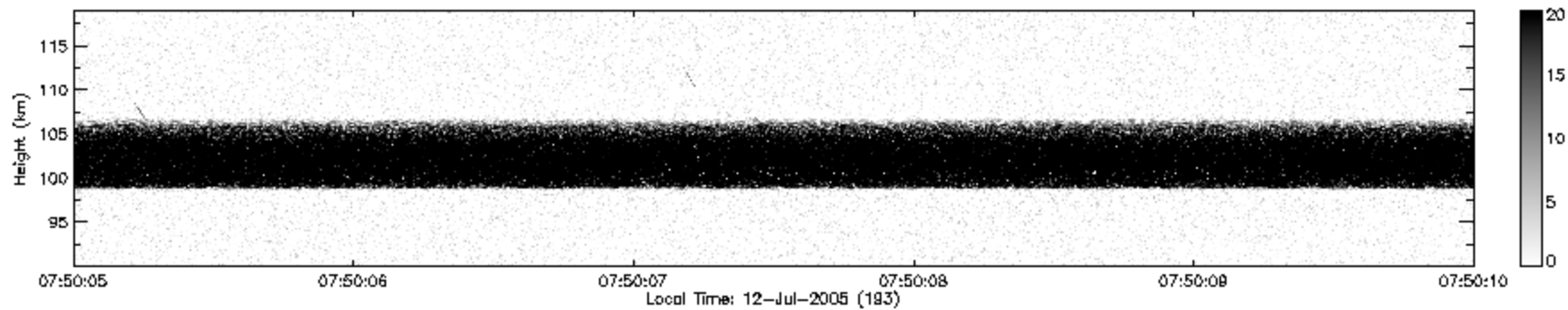
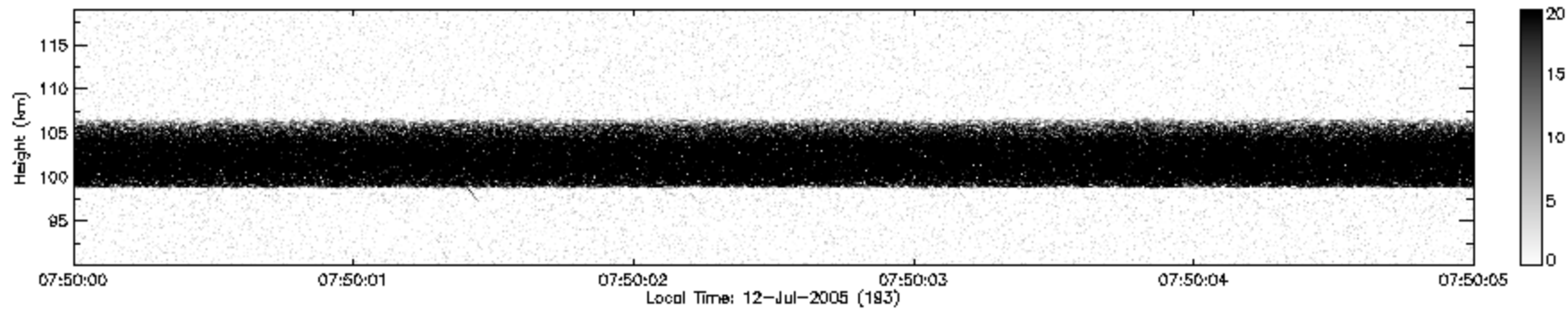
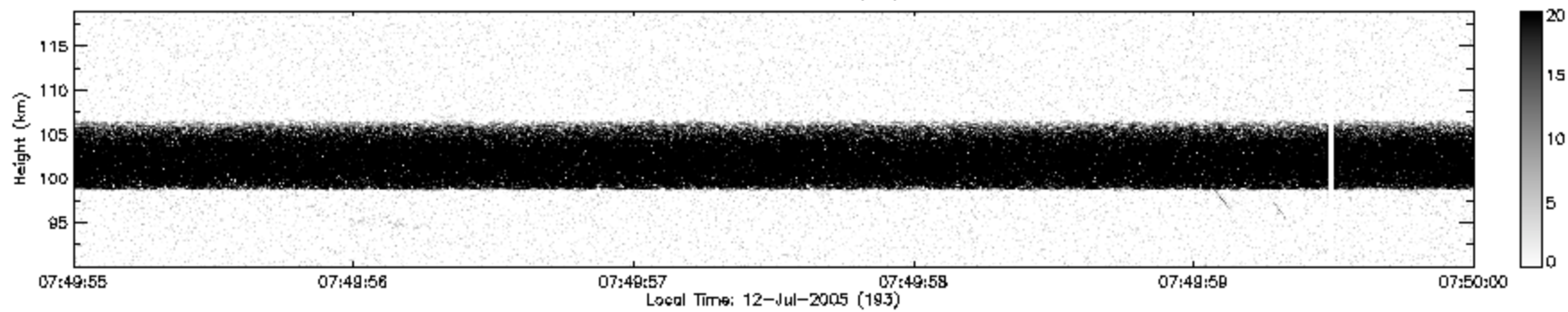
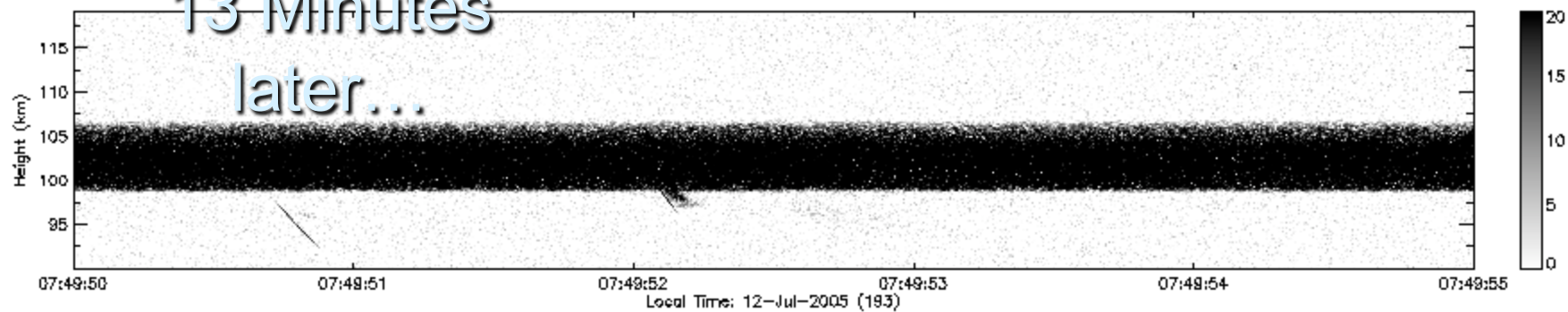


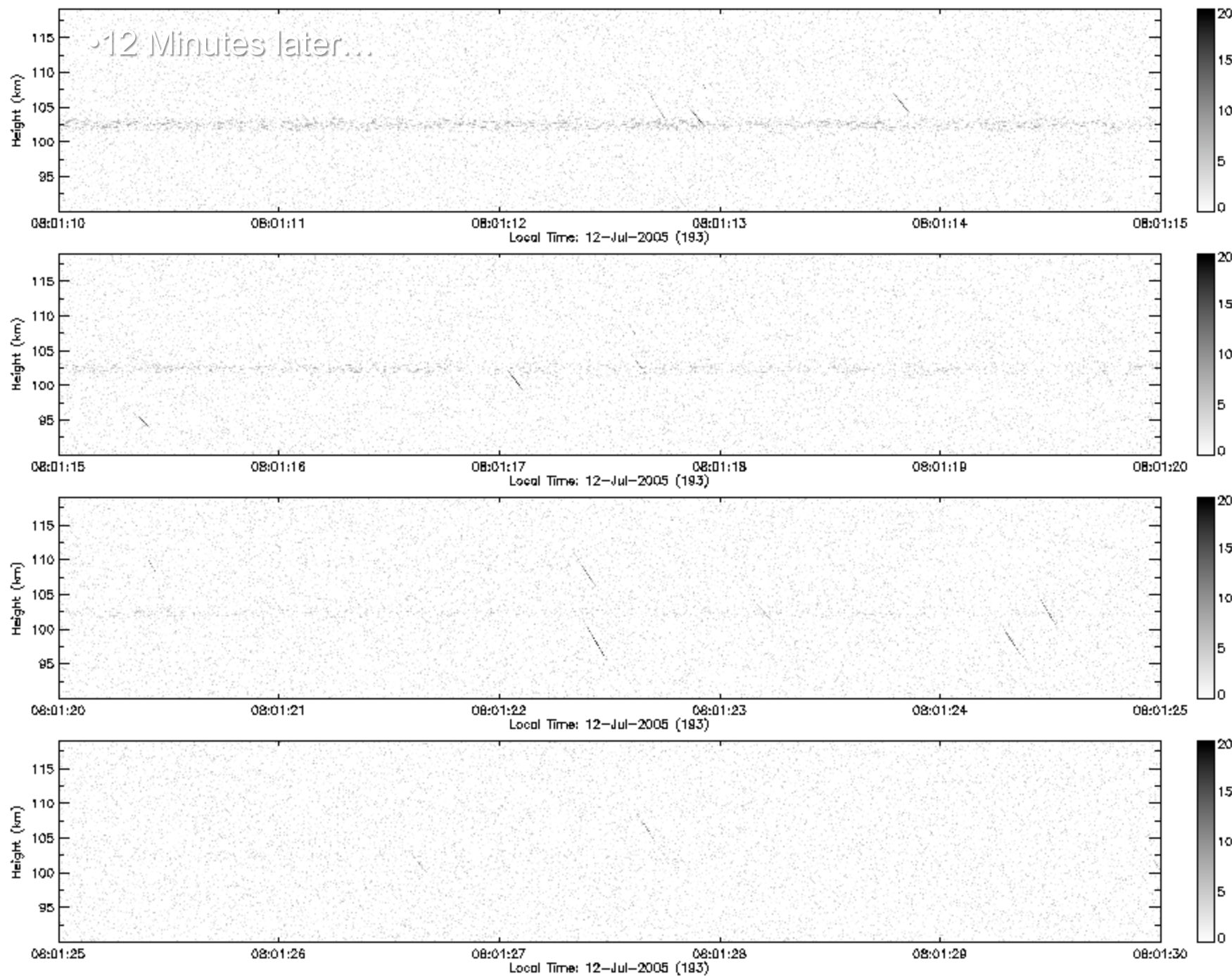
Daytime...





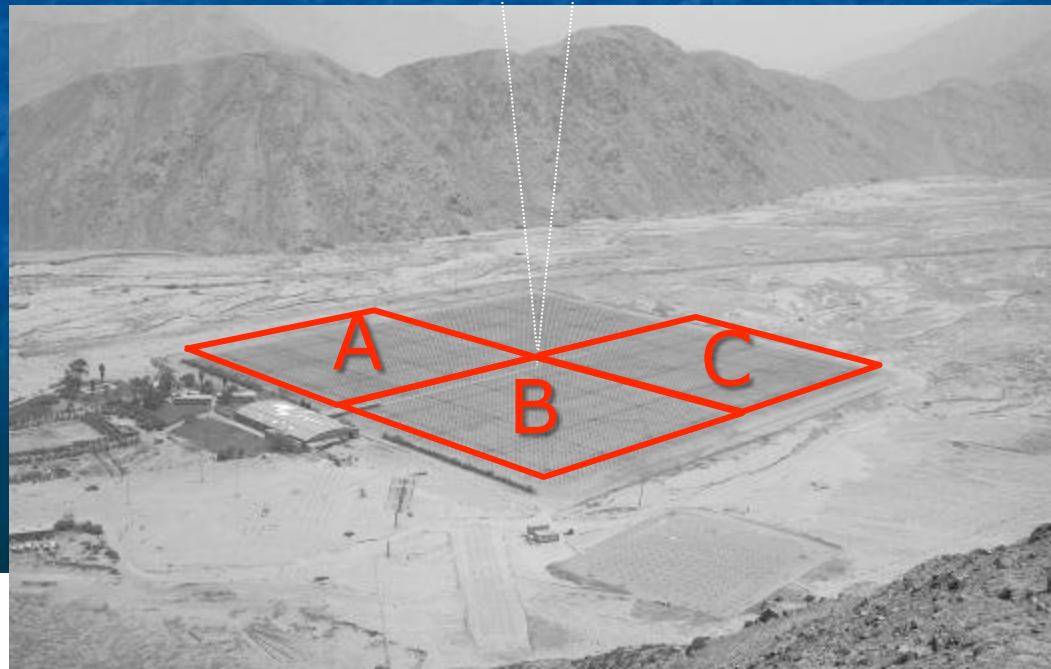
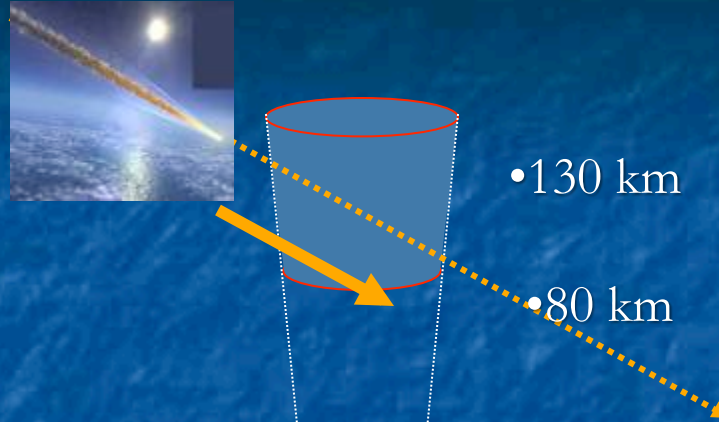
13 Minutes
later...





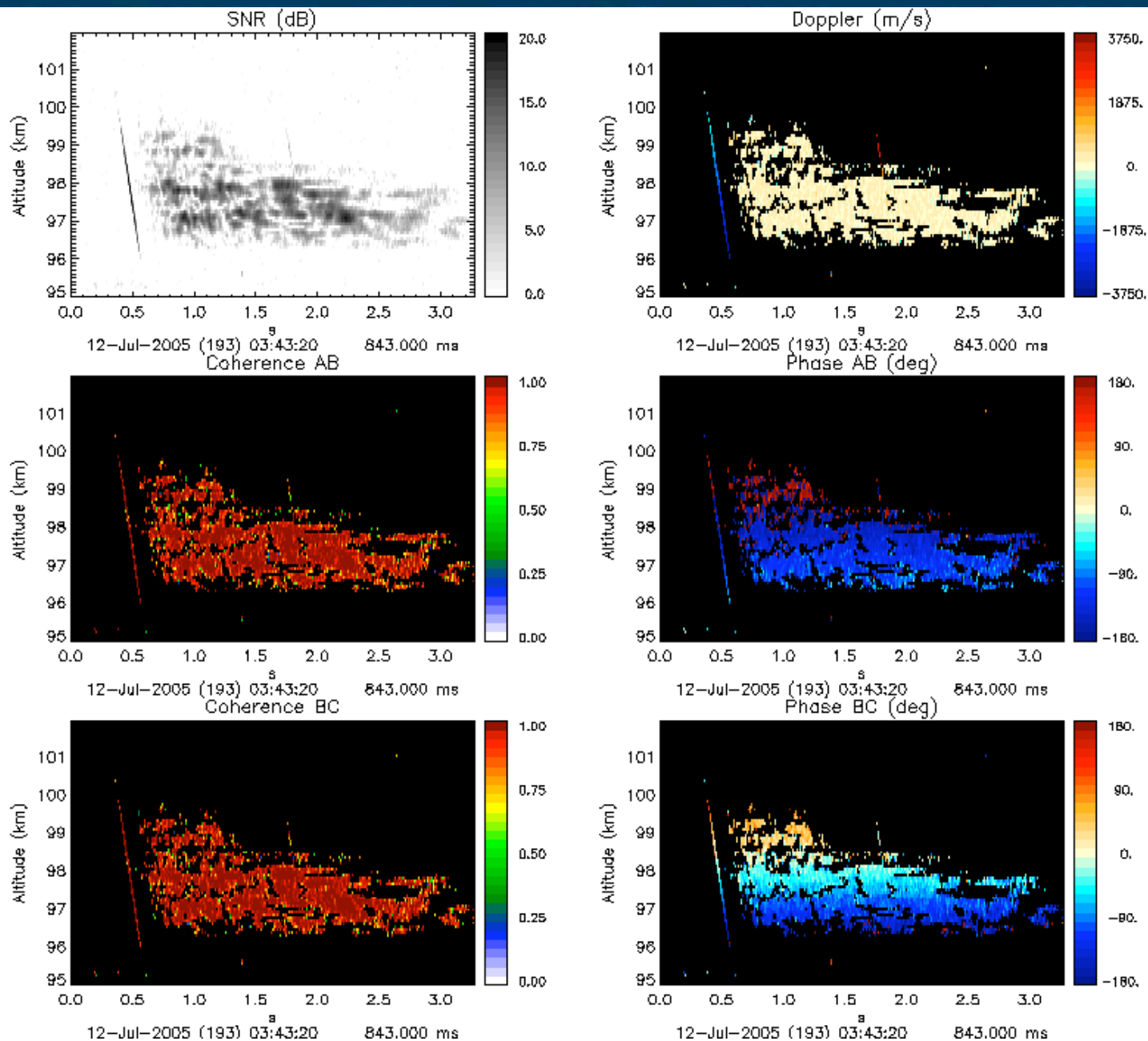
JRO 50MHz Radar Observations

- Antenna:
 - 300m x 300m
 - 18,432 dipoles
 - Peak power: 2 MW
 - Frequency: 50 MHz
- *A truly High-Power Large-Aperture Radar (HPLA)*
- **Interferometer**

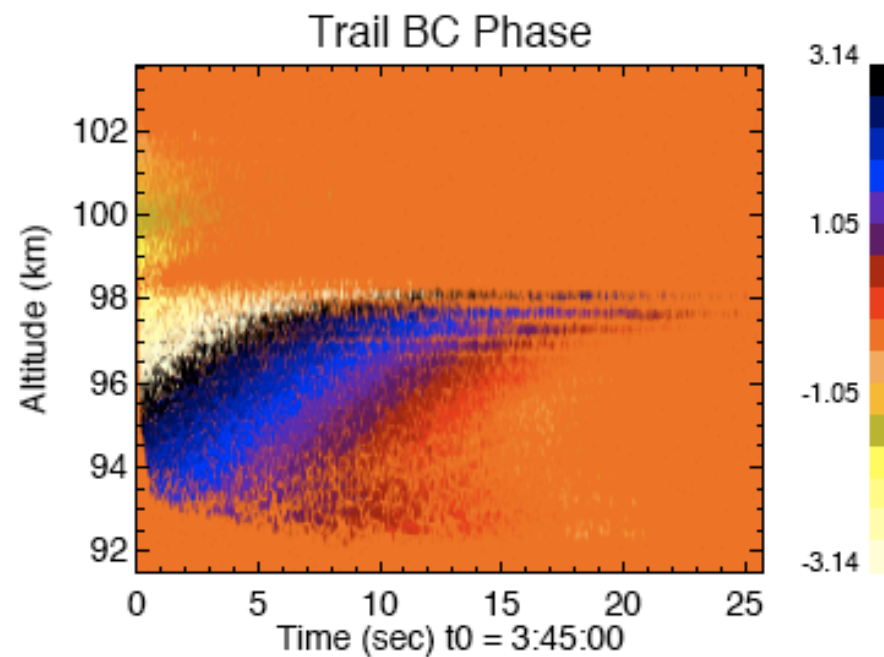
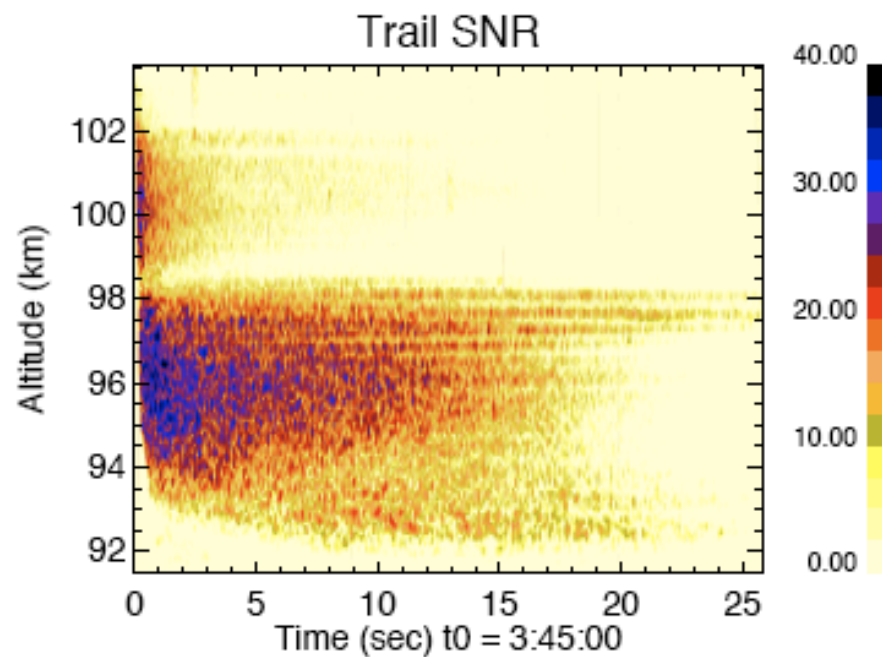
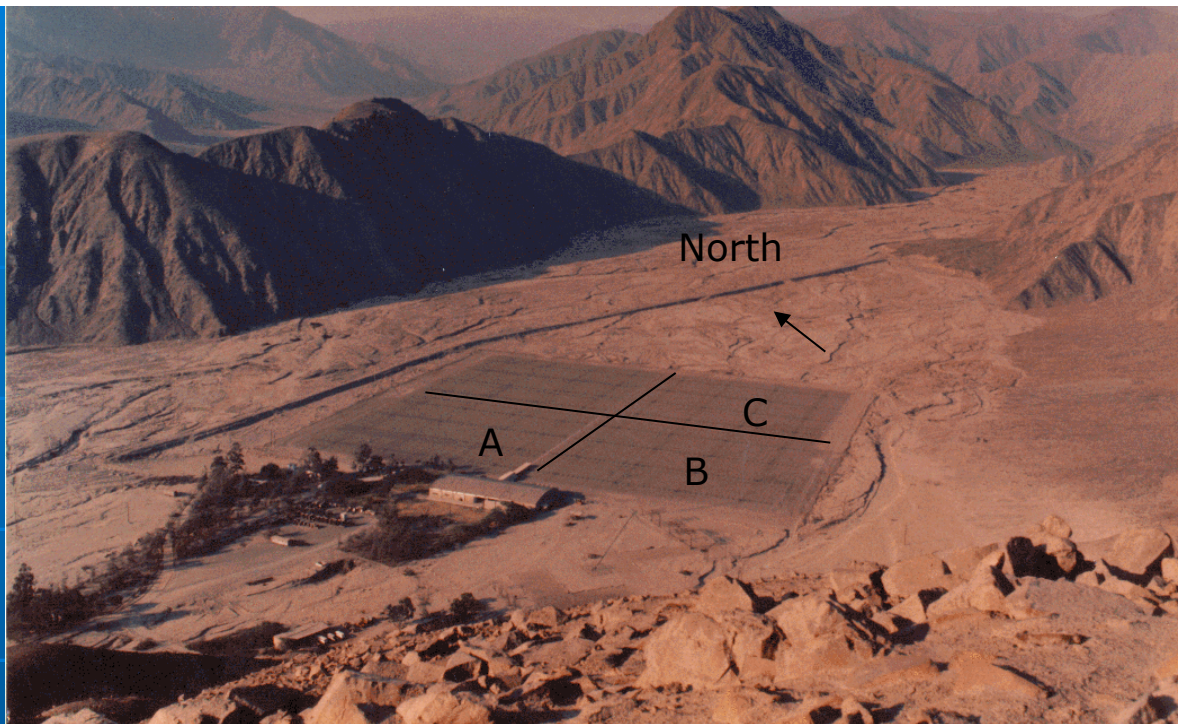


•Not to scale

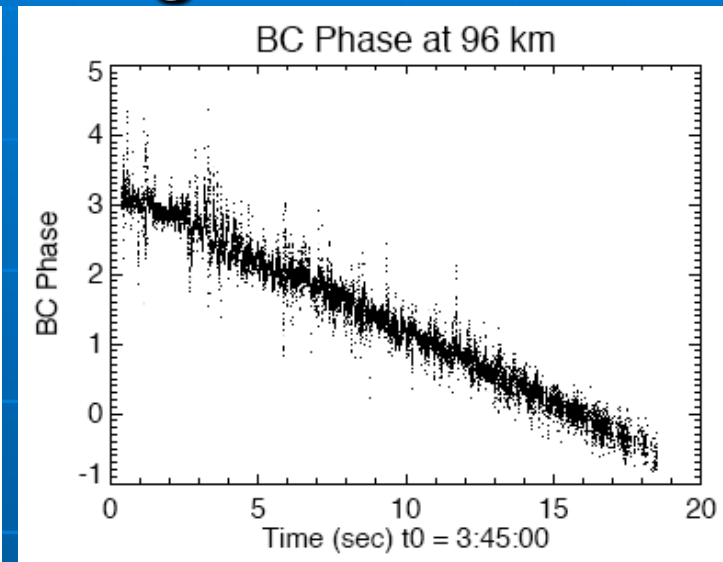
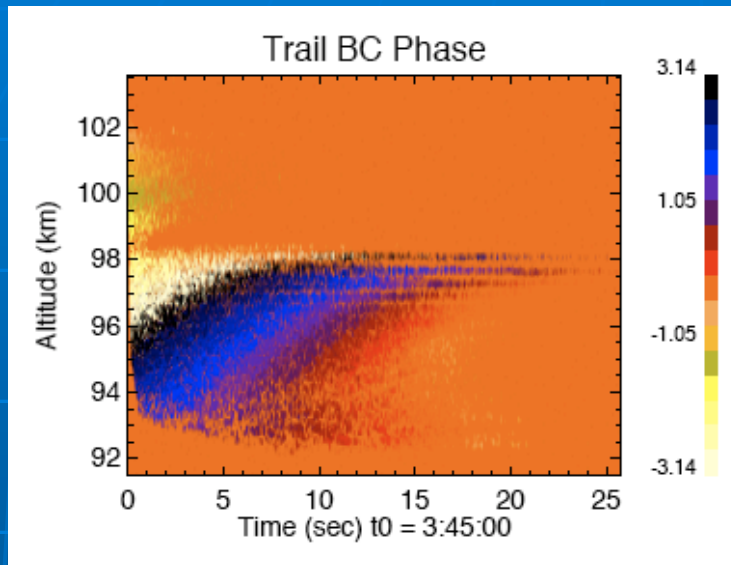
SNR, Doppler, and Phase



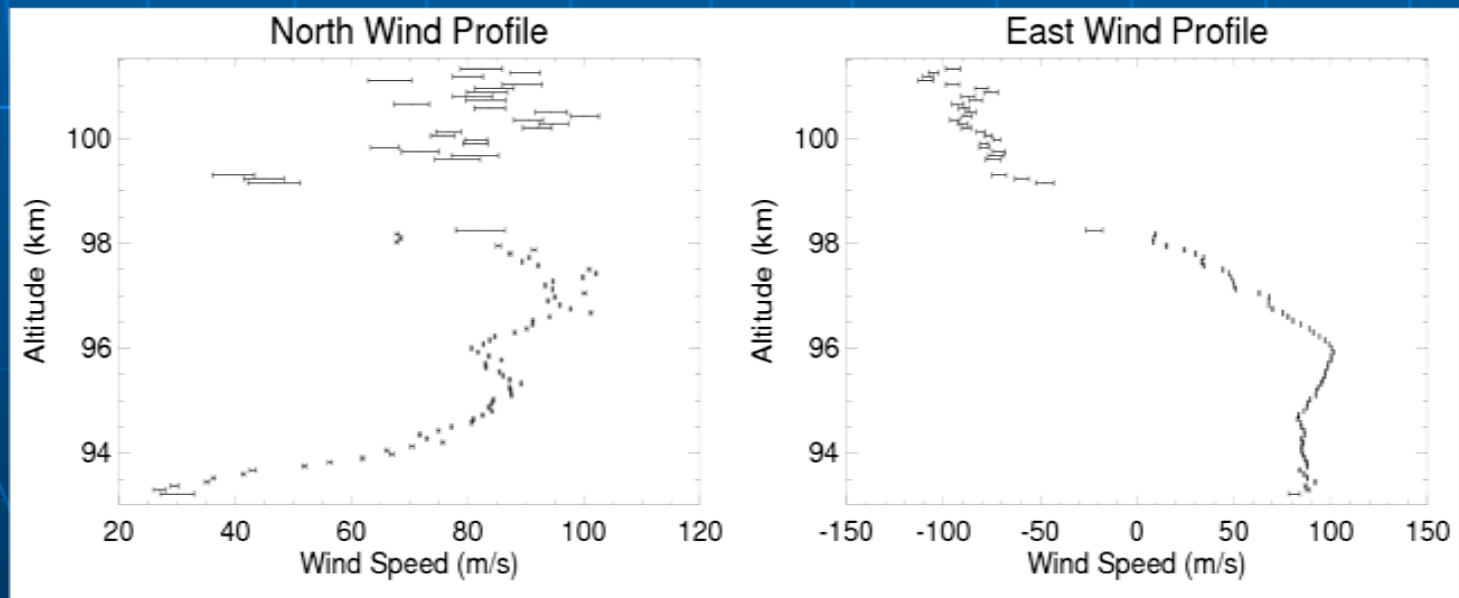
Jicamarca Trail Interferometry



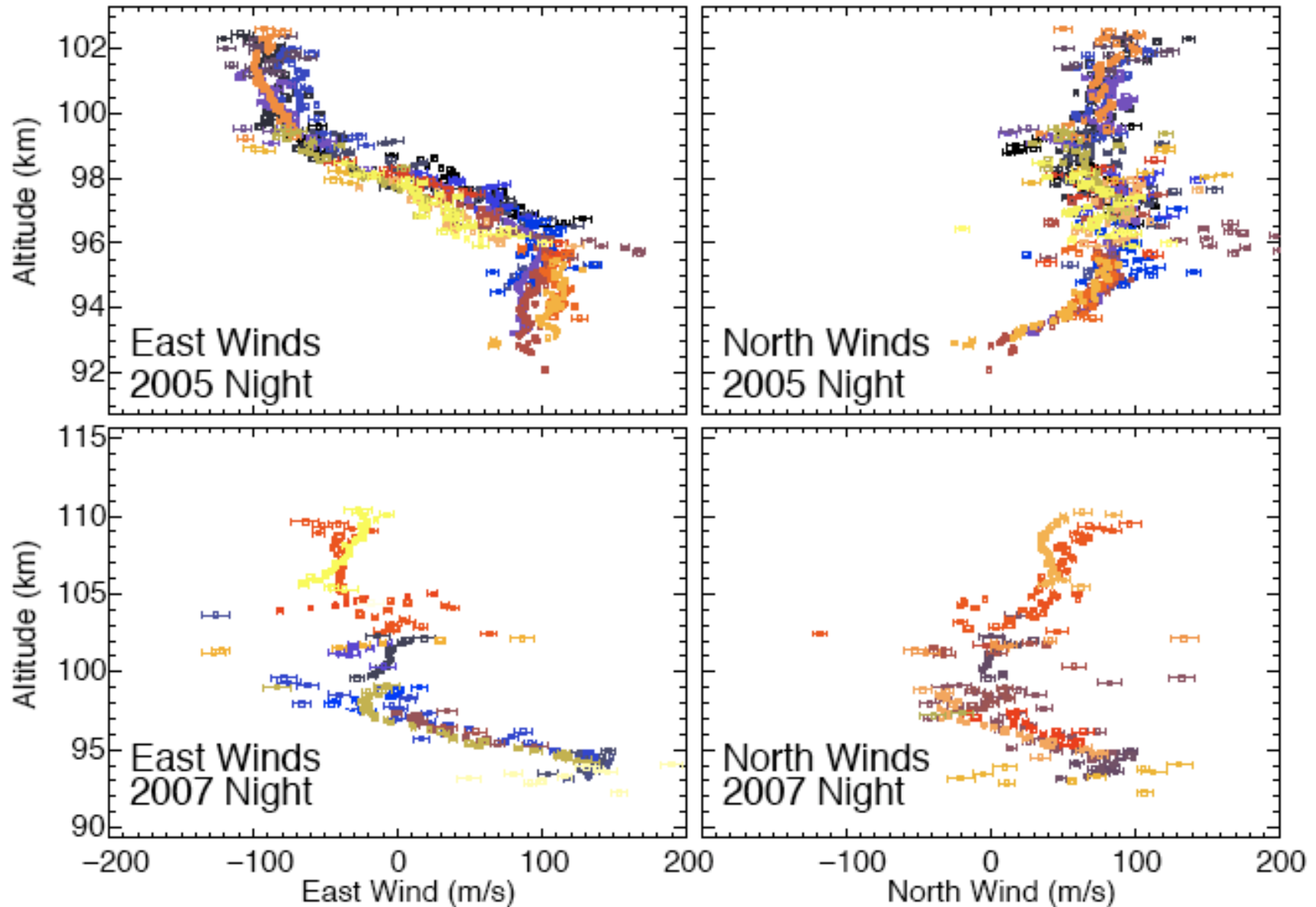
Trail Phase at Single Altitude



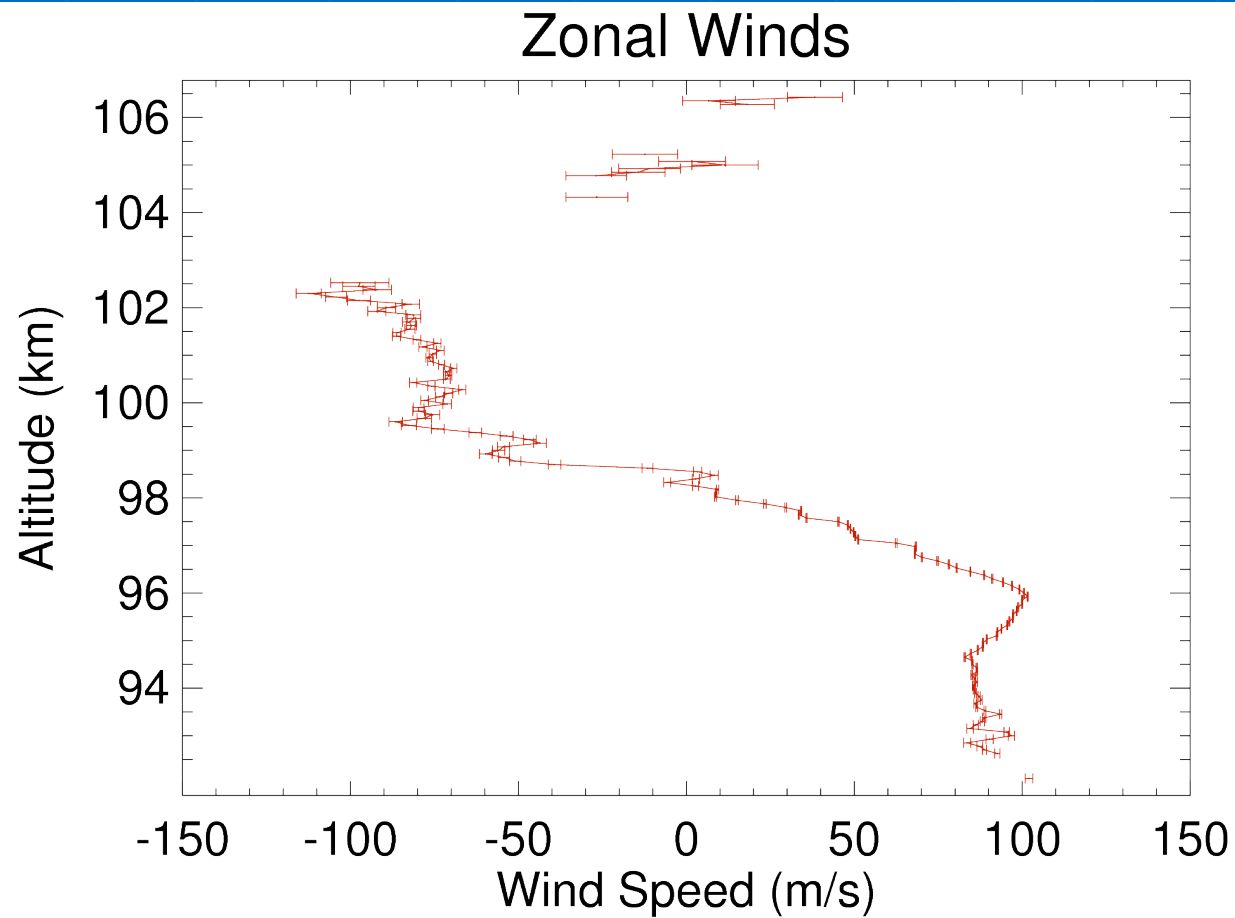
Winds from all altitudes



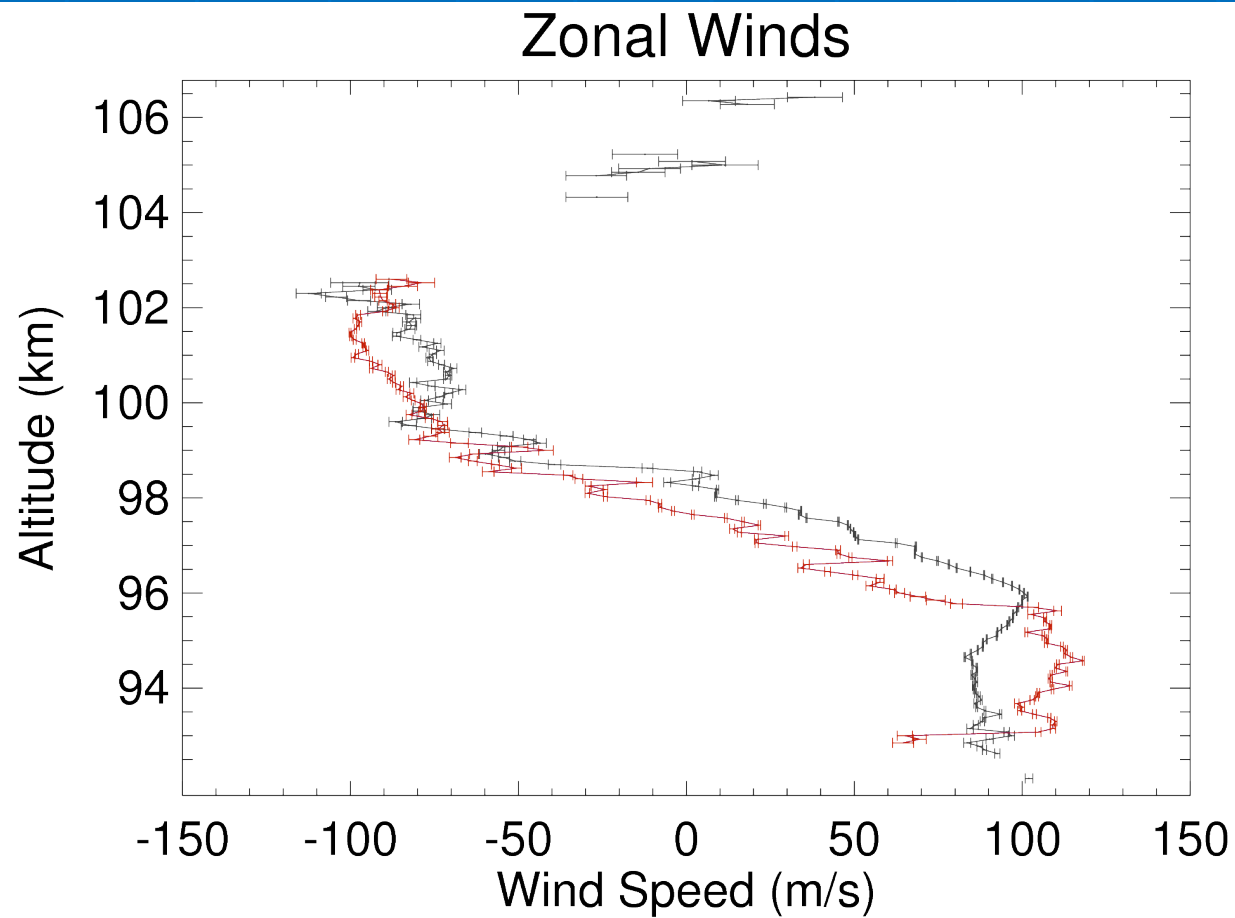
Wind Profiles from many trails



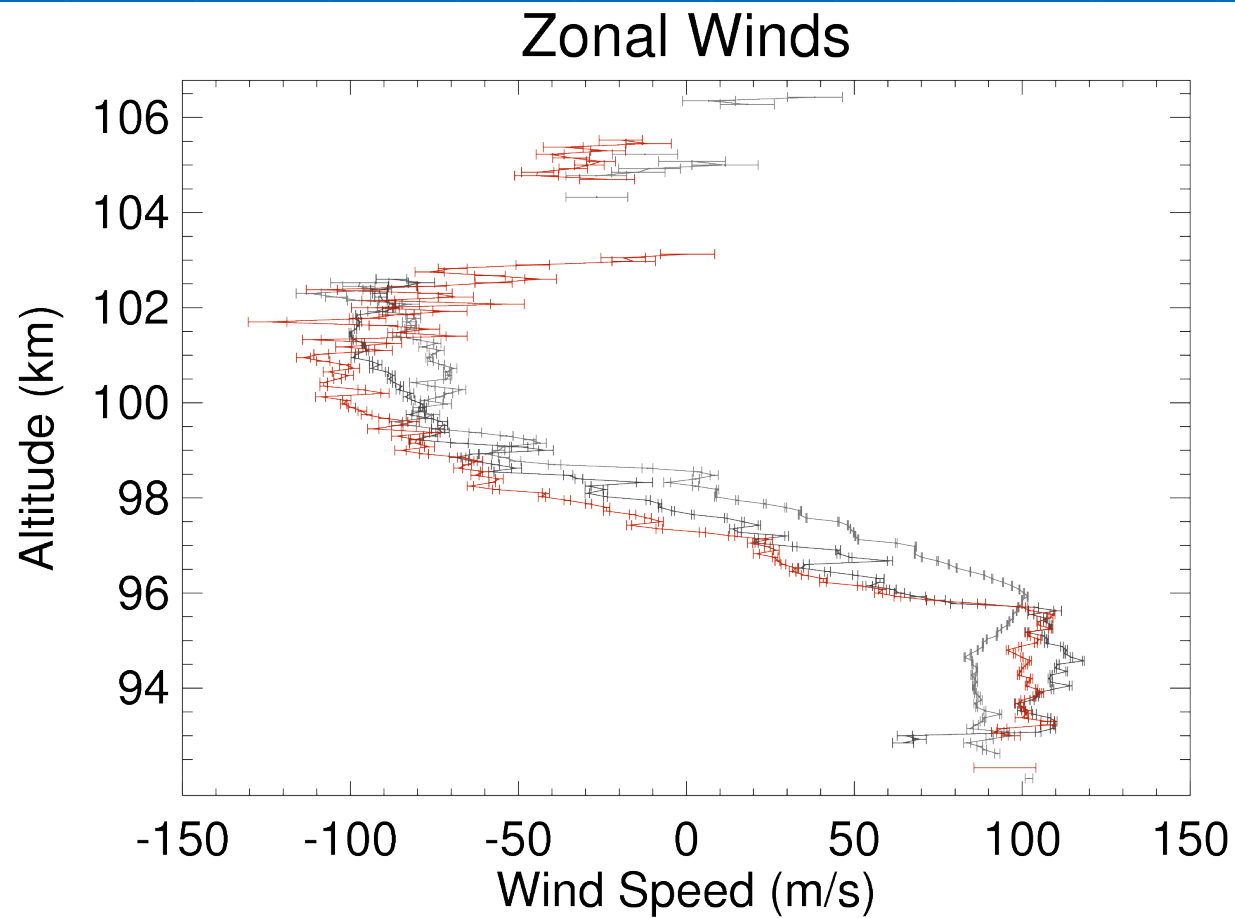
Zonal Winds from 15 min of data



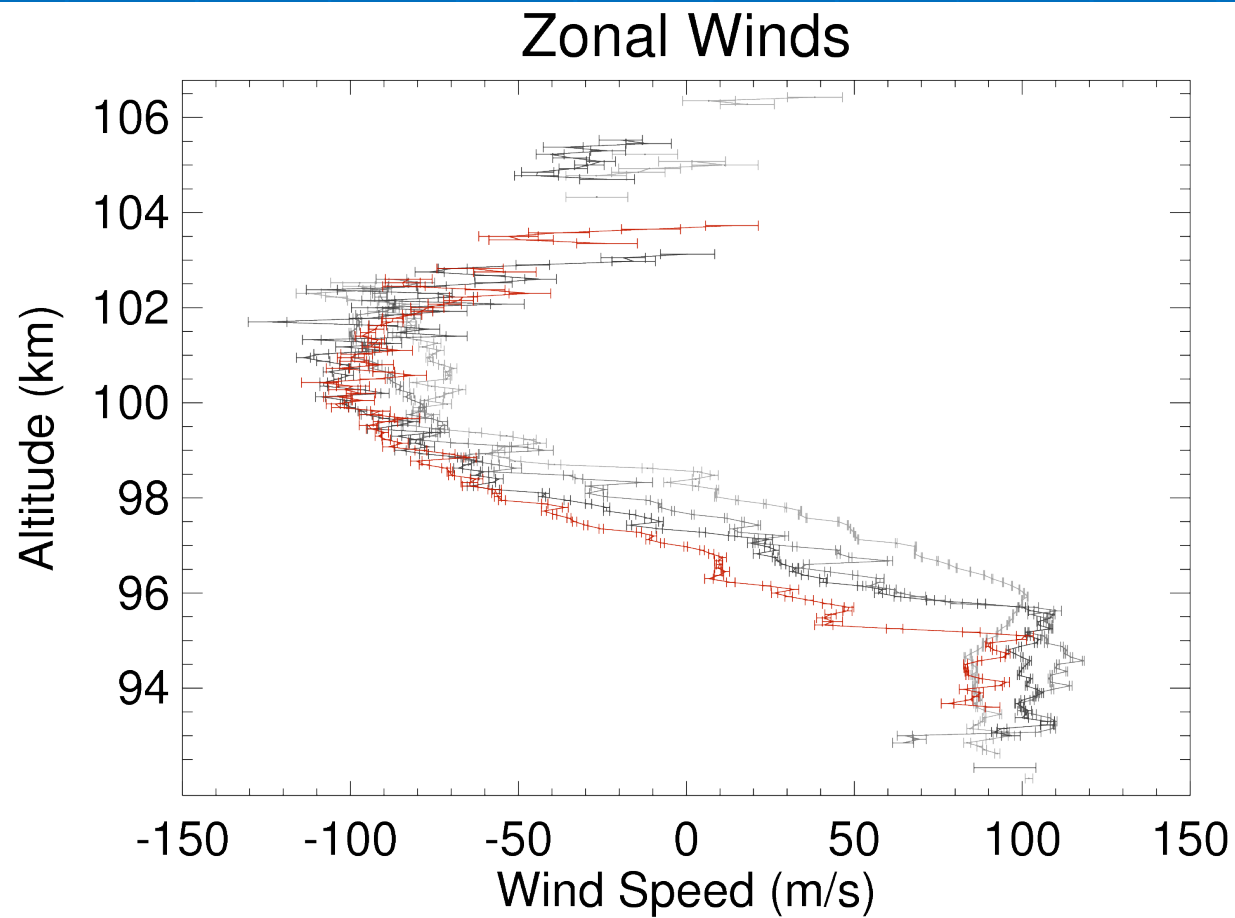
Zonal Winds from 15 min of data



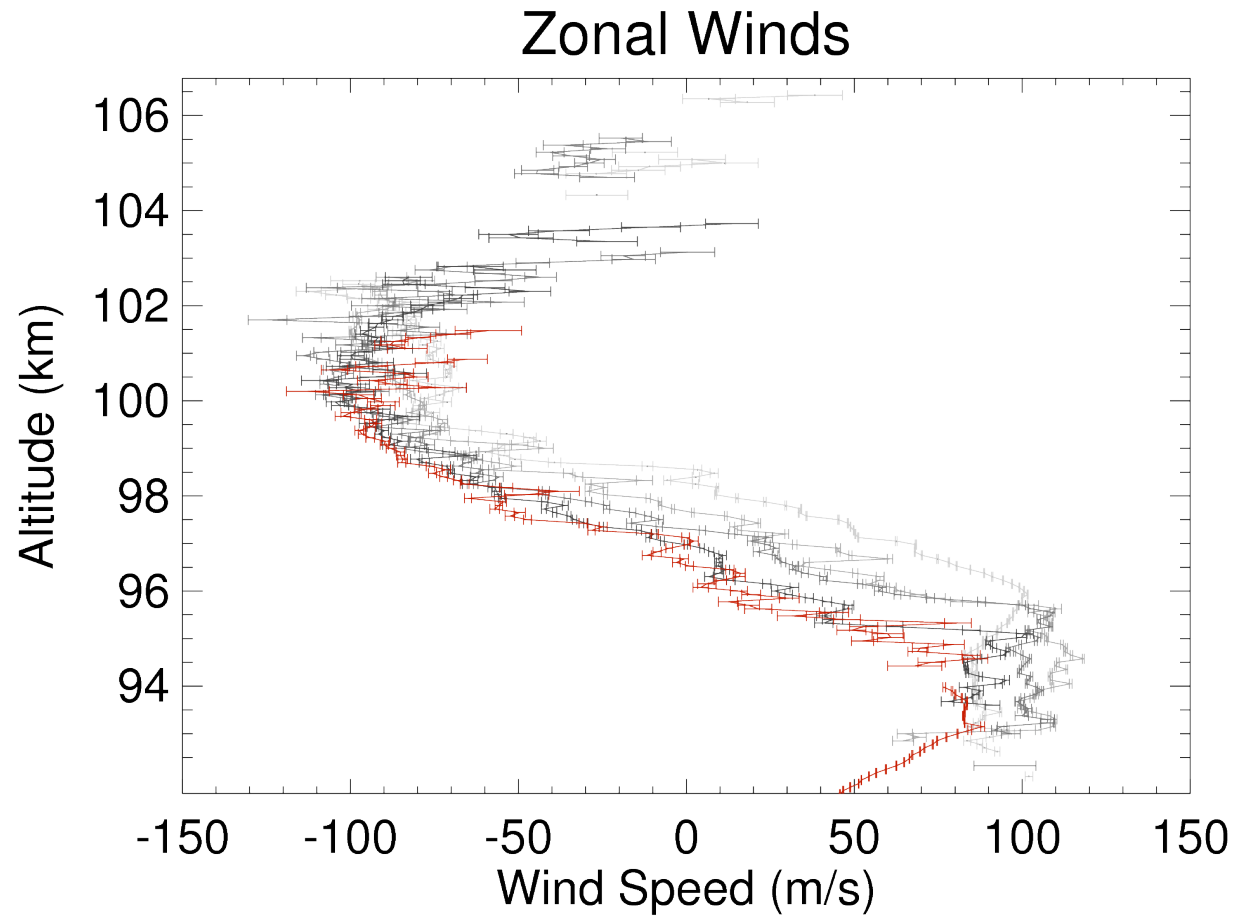
Zonal Winds from 15 min of data



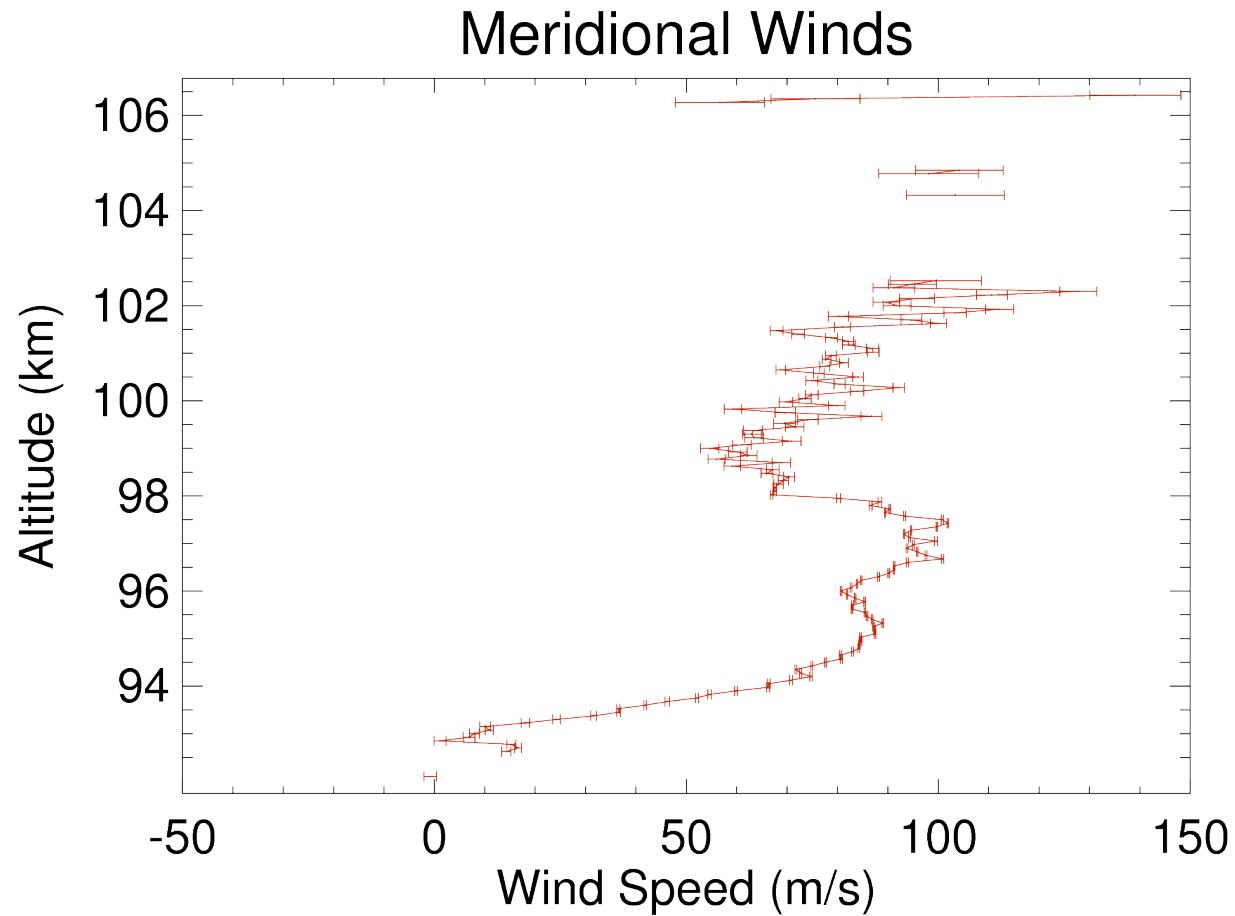
Zonal Winds from 15 min of data



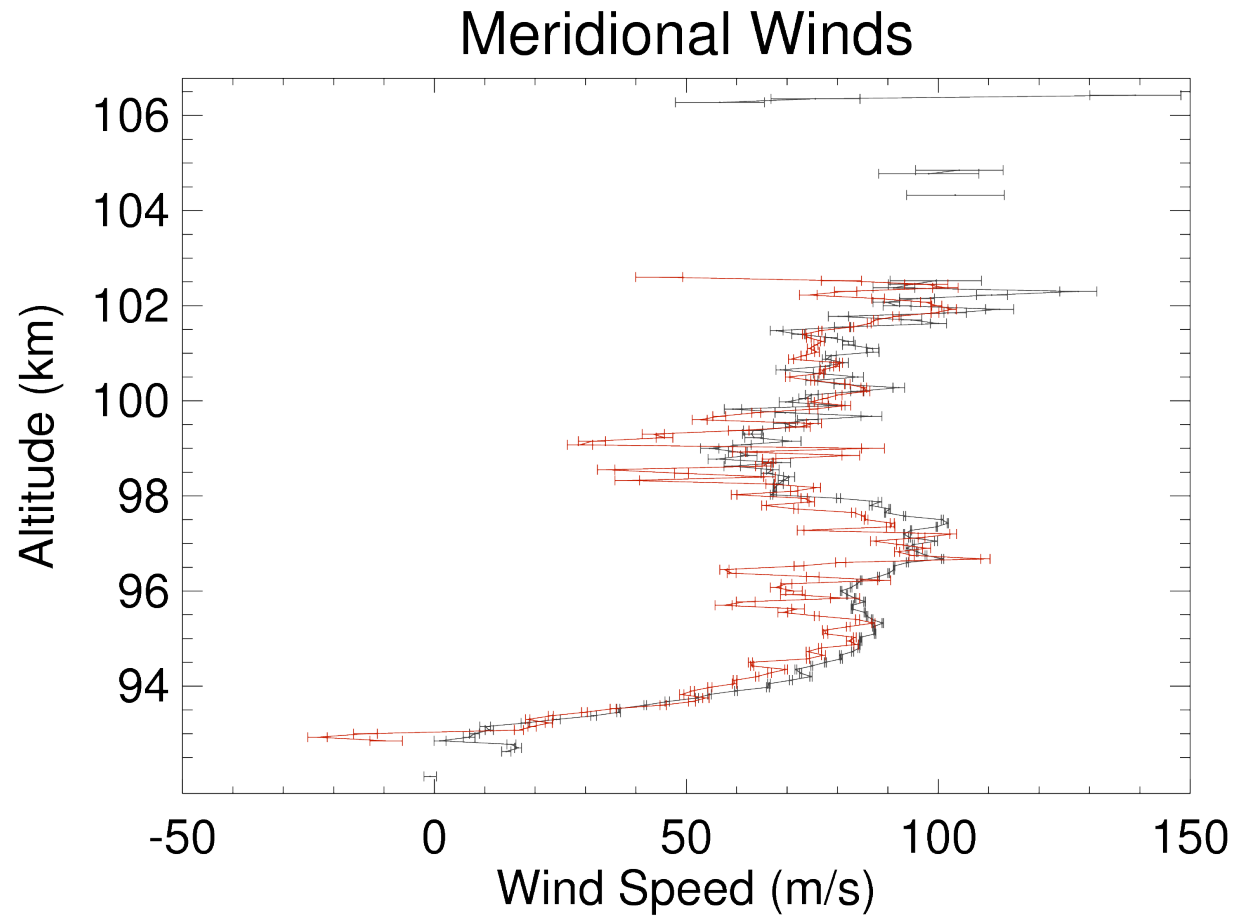
Zonal Winds from 15 min of data



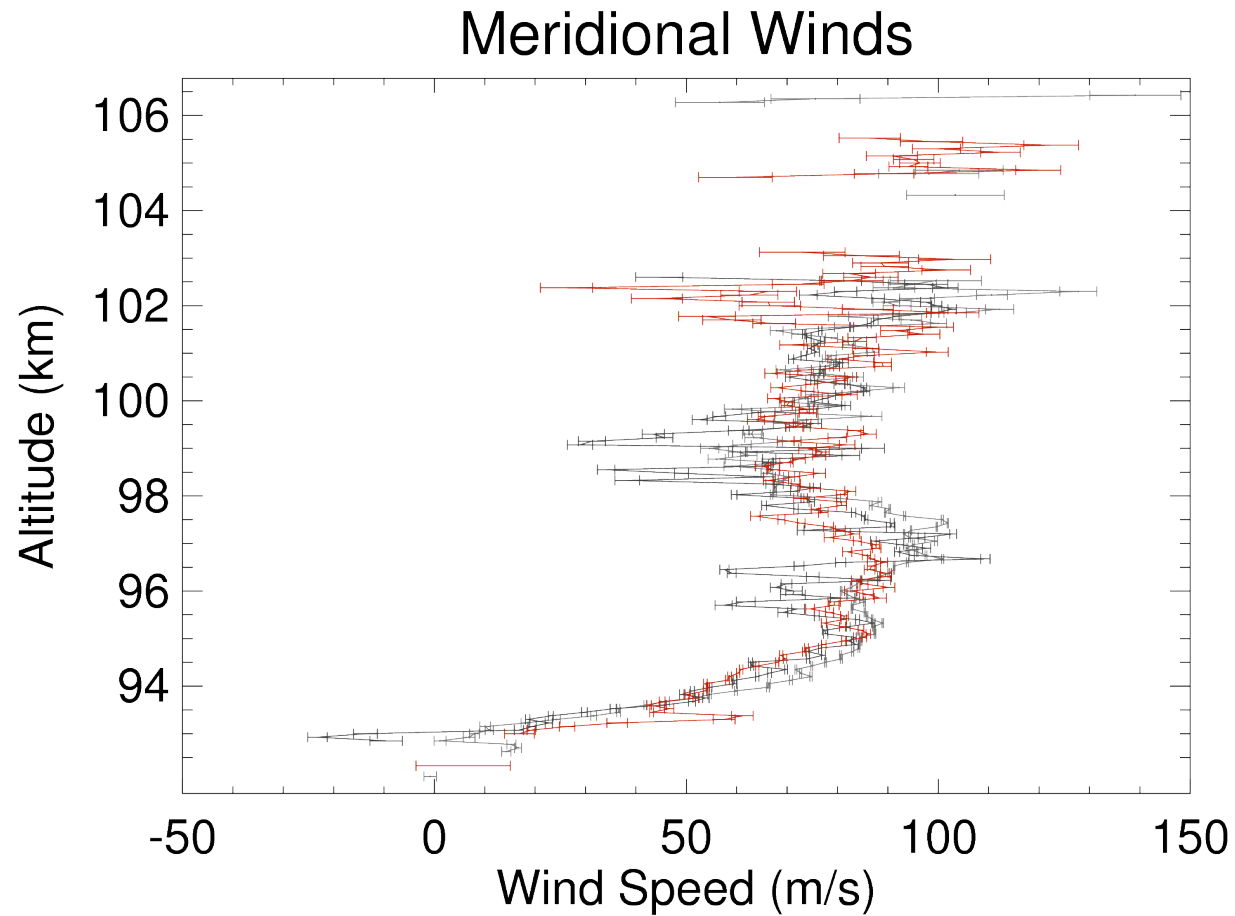
Meridional Winds from 15 min of data



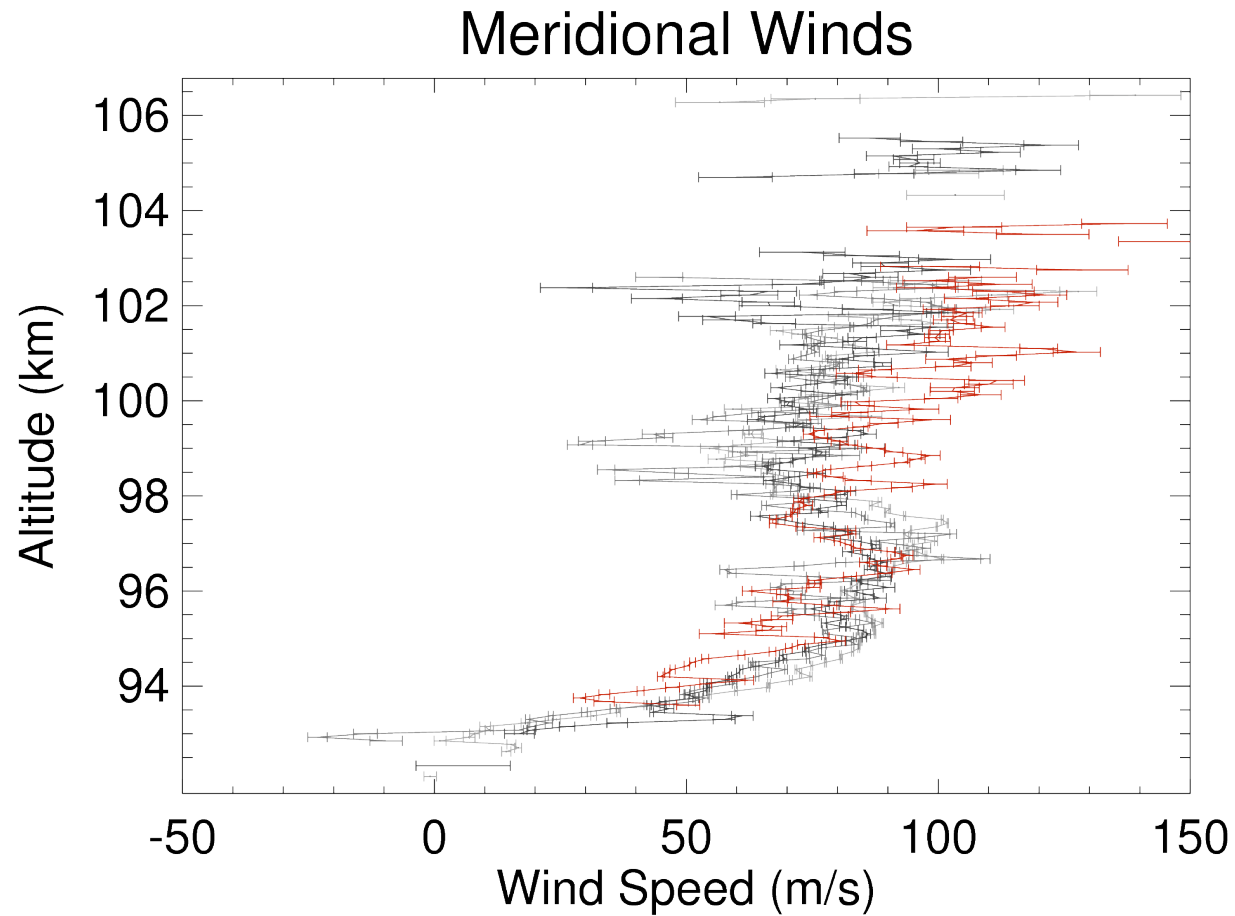
Meridional Winds from 15 min of data



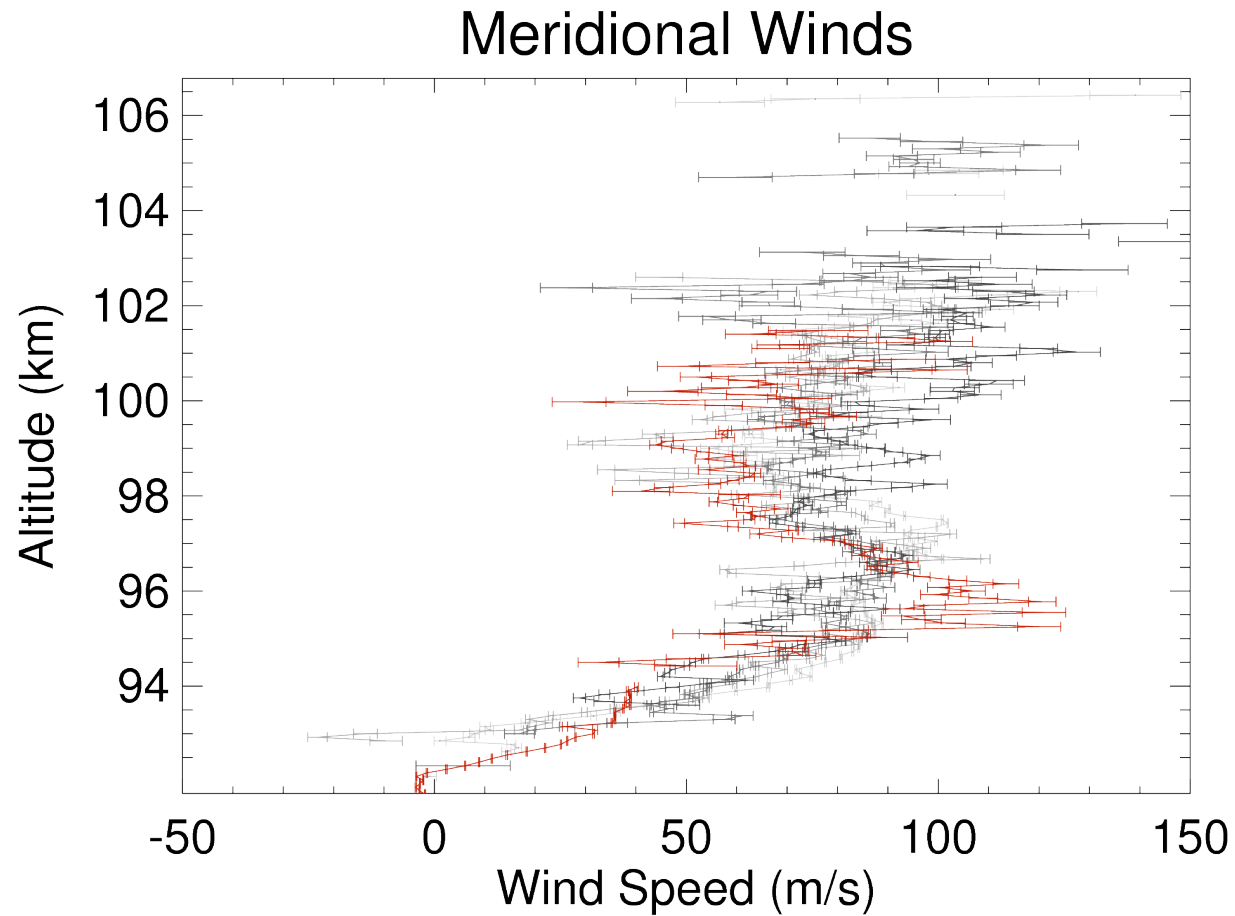
Meridional Winds from 15 min of data



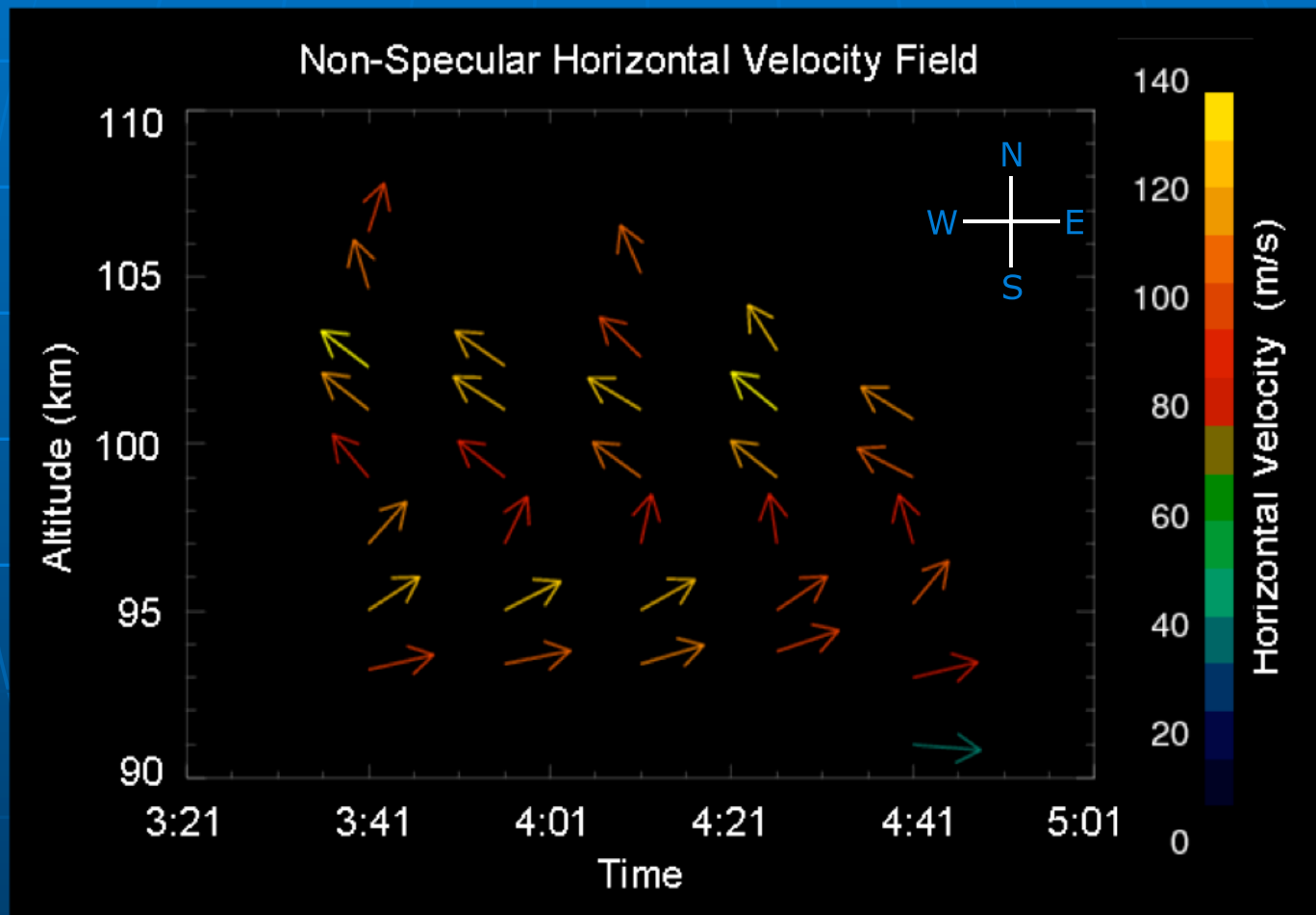
Meridional Winds from 15 min of data



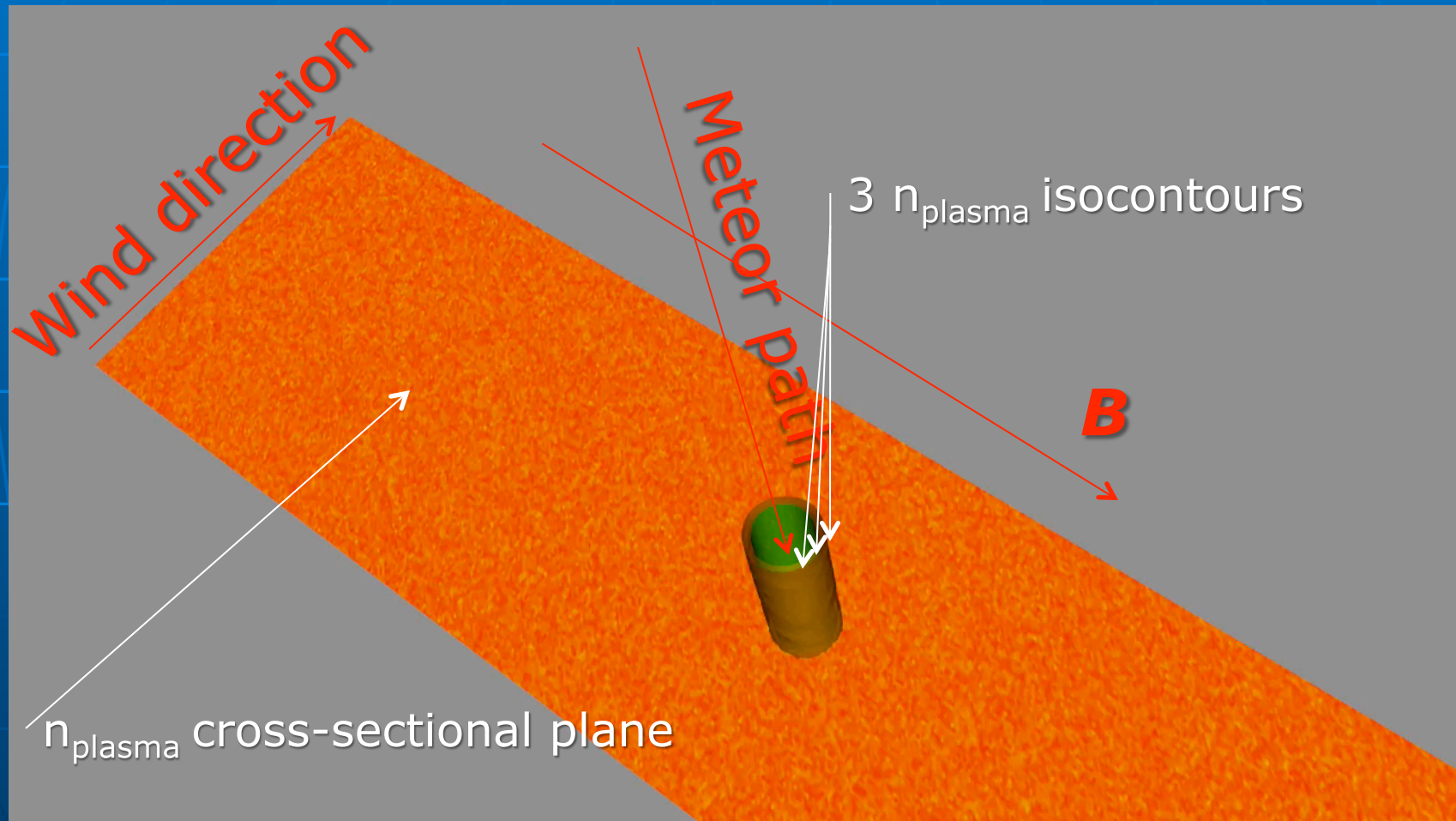
Meridional Winds from 15 min of data



Wind fields

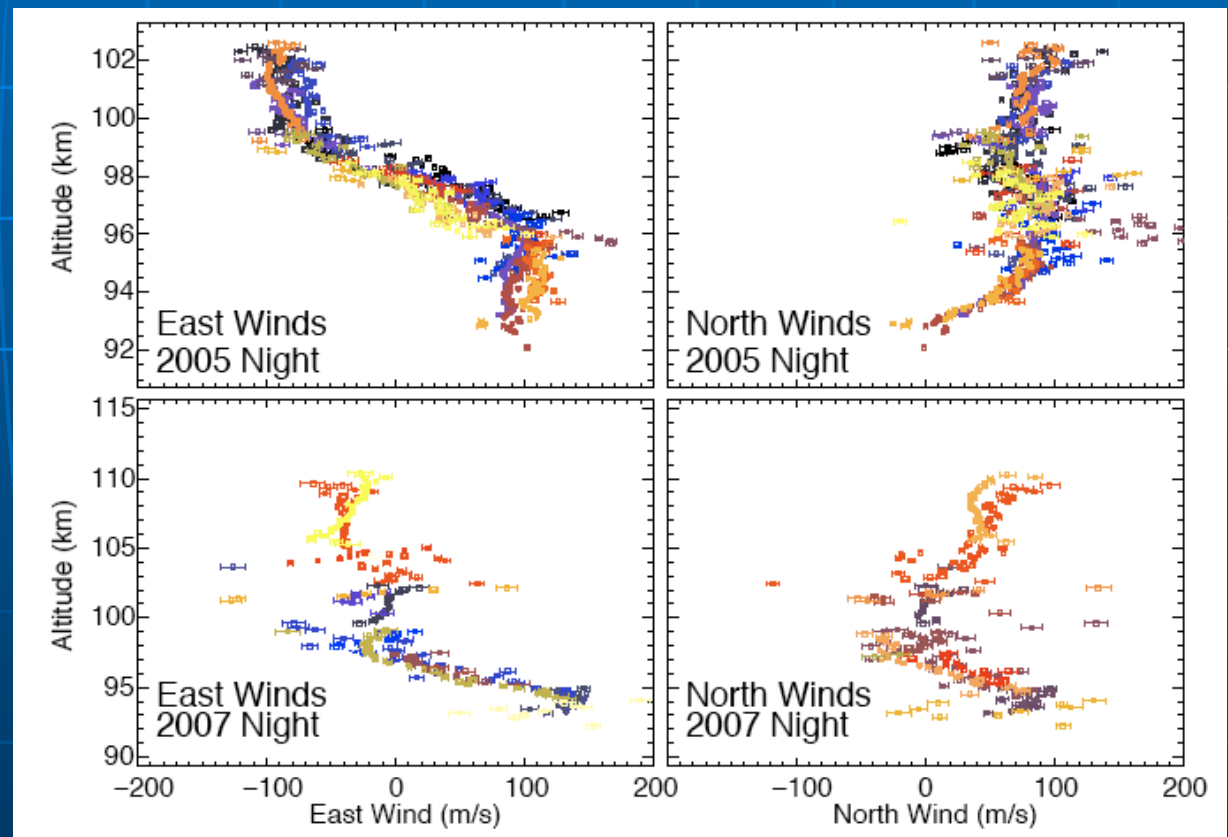
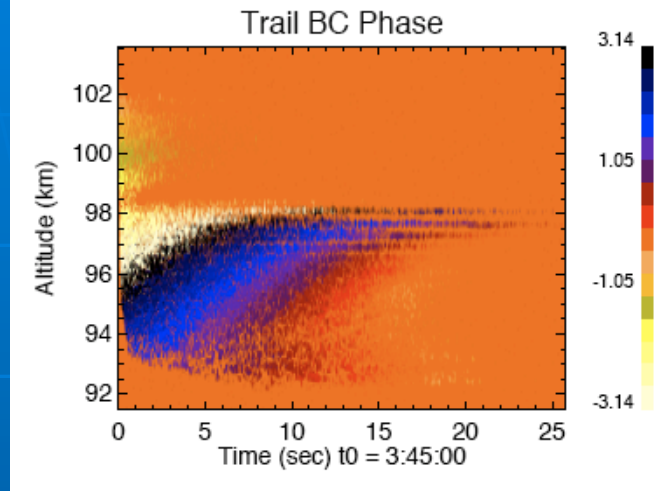


Particle in Cell Simulations of Meteor Plasma



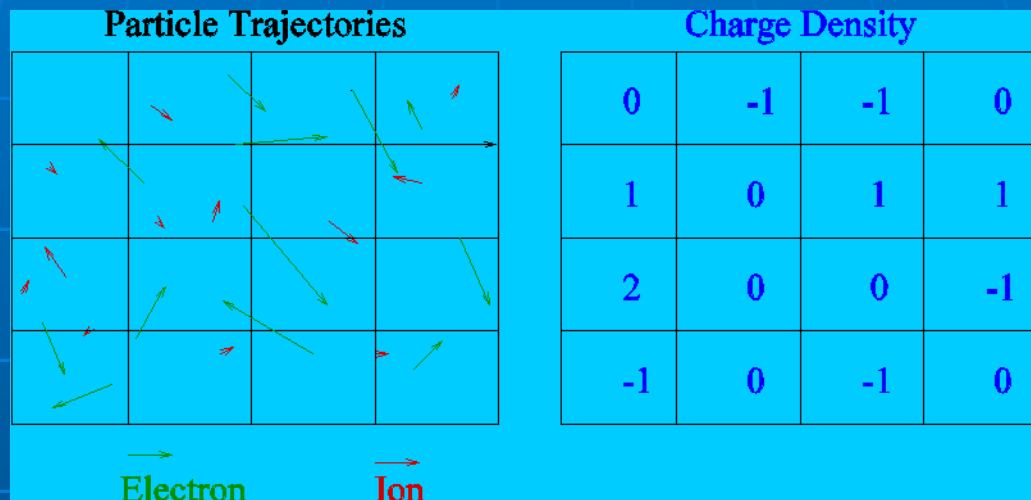
Conclusions

- New Tool to Monitor Winds
- High Spatial and Time Resolution
- Low and Mid-latitudes
- Publications:
 - GRL, May '09
 - Physics Today, June '09



Simulation Method: Particle-In-Cell

1. Gather to determine charge density, ρ



2. Calculate Electric field:
3. Update velocities:
4. Update Positions:
5. Collide particles with neutrals
6. Go to Step 1

$$\vec{\nabla} \cdot \vec{E} = \rho / \epsilon_0$$

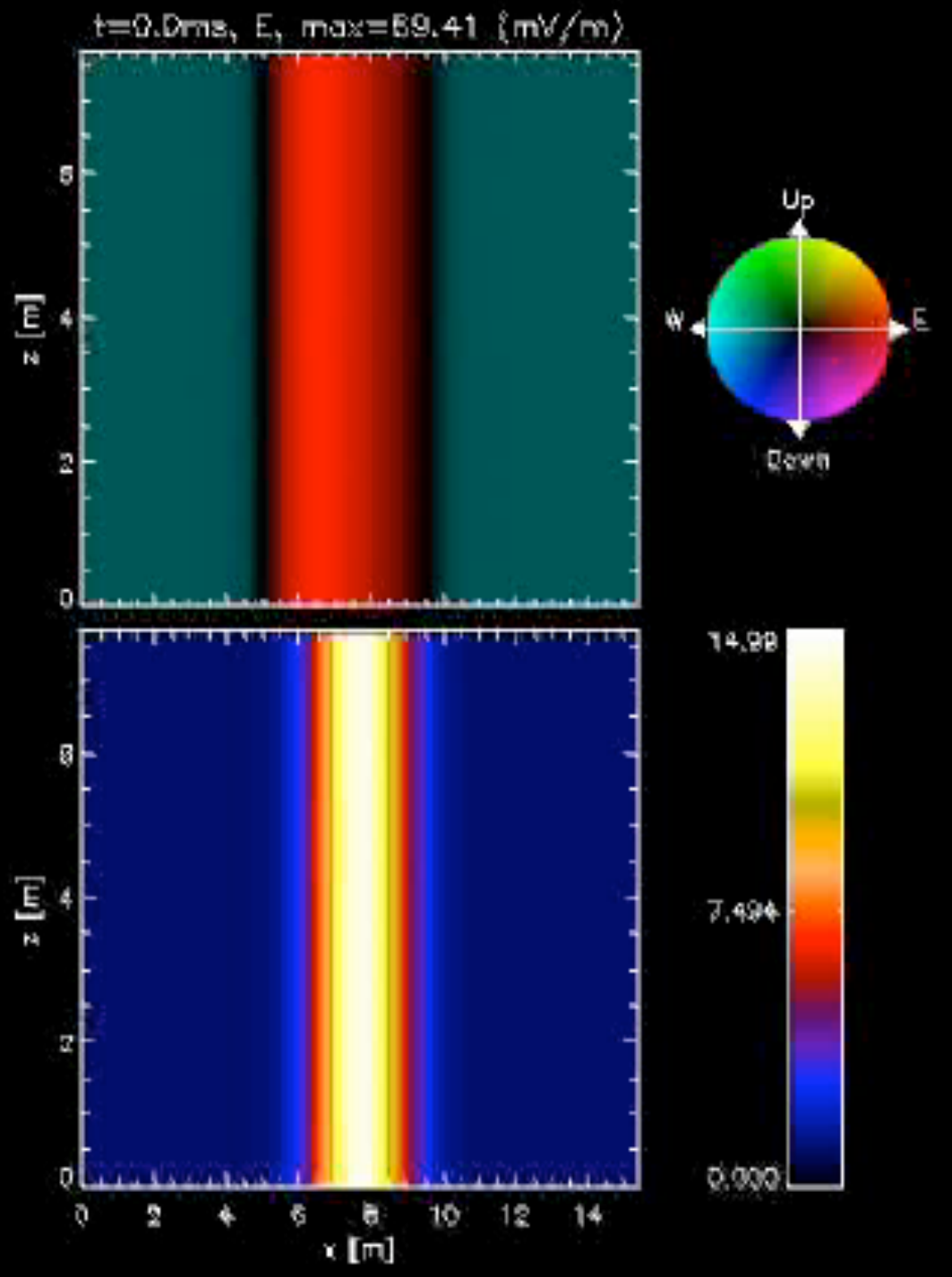
$$\frac{d\vec{v}_i}{dt} = \frac{q_i}{m_i} [E(\vec{x}_i, t) + \vec{v}_i \times B(\vec{x}_i, t)]$$

$$\frac{d\vec{x}_i}{dt} = \vec{v}_i$$

Meteor Plasma Simulation in 2D

Electric Field

Density



3D Meteor Simulations

