

Mass Distributions Found Using the Jicamarca Radio Observatory

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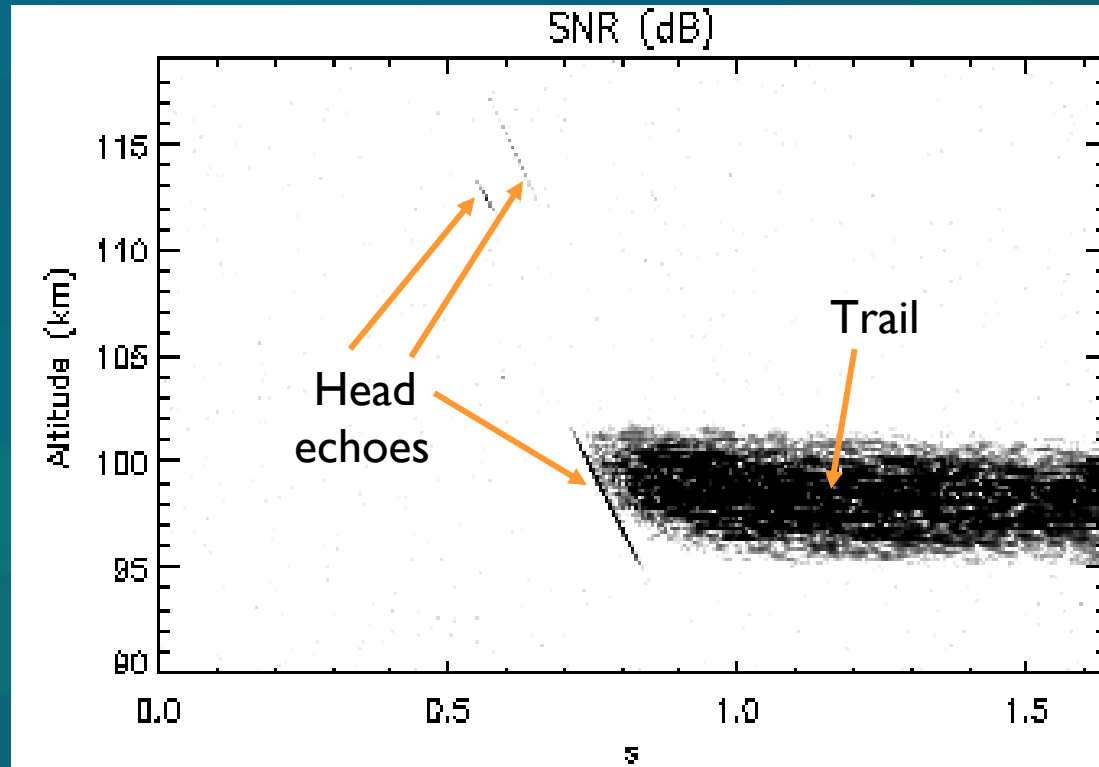
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Jicamarca Radio Observatory

Outline

- Introduction
- Observations
- Mass Distribution
- Spatial Distributions
- Mass Index
- Conclusions

Introduction



Introduction - calculating mass

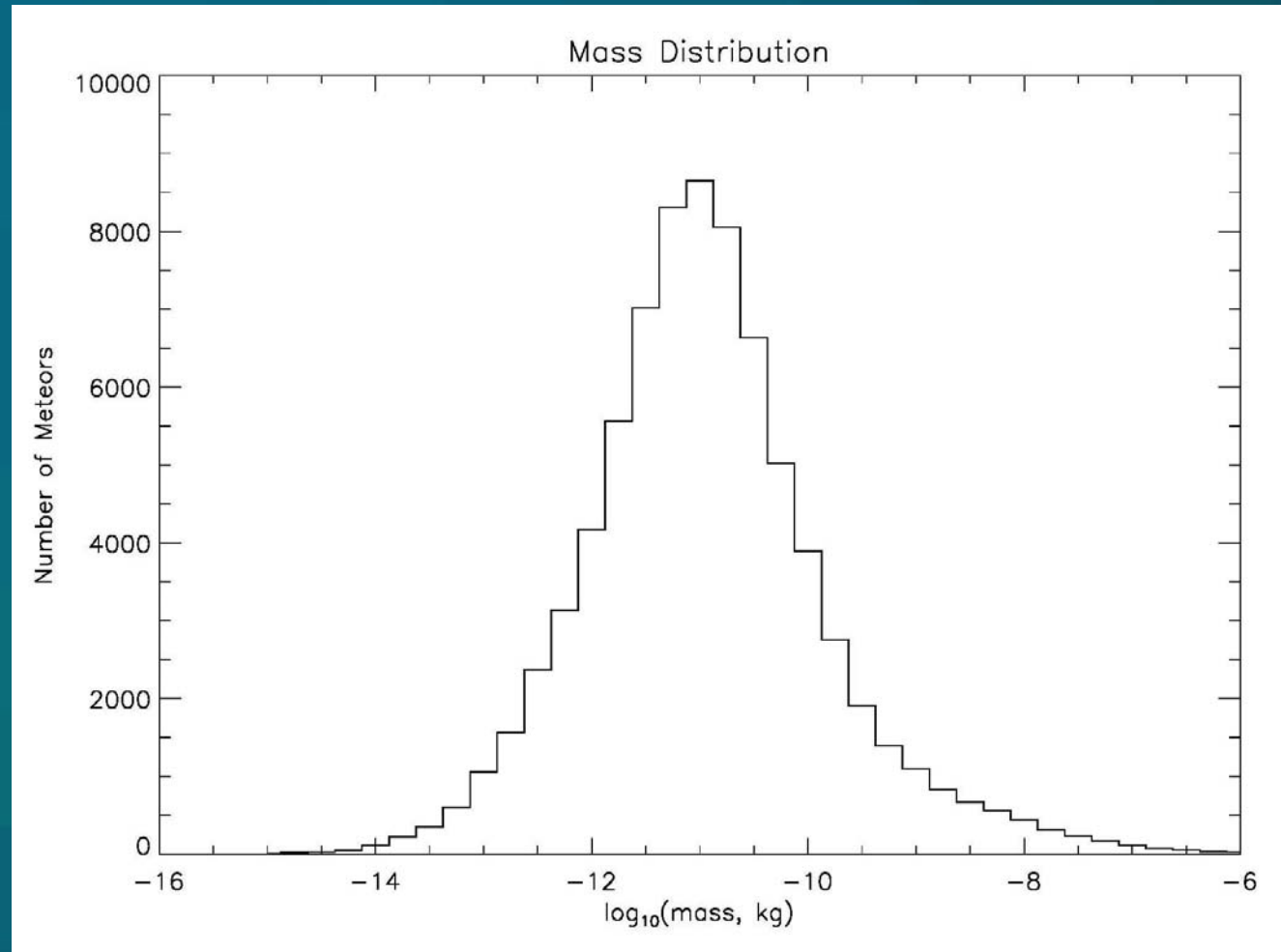
- Number of different methods
 - dynamical mass
 - scattering mass (Close et al. 2005)
 - ablation model (Bass et. al 2008)
 - differential ablation model (Janches et al. 2009)
- Dynamical mass:

$$m = \frac{9\pi}{2\rho_m^2} \left(\frac{v^2 \gamma \rho_{air}}{dv/dt} \right)^3$$

Observations

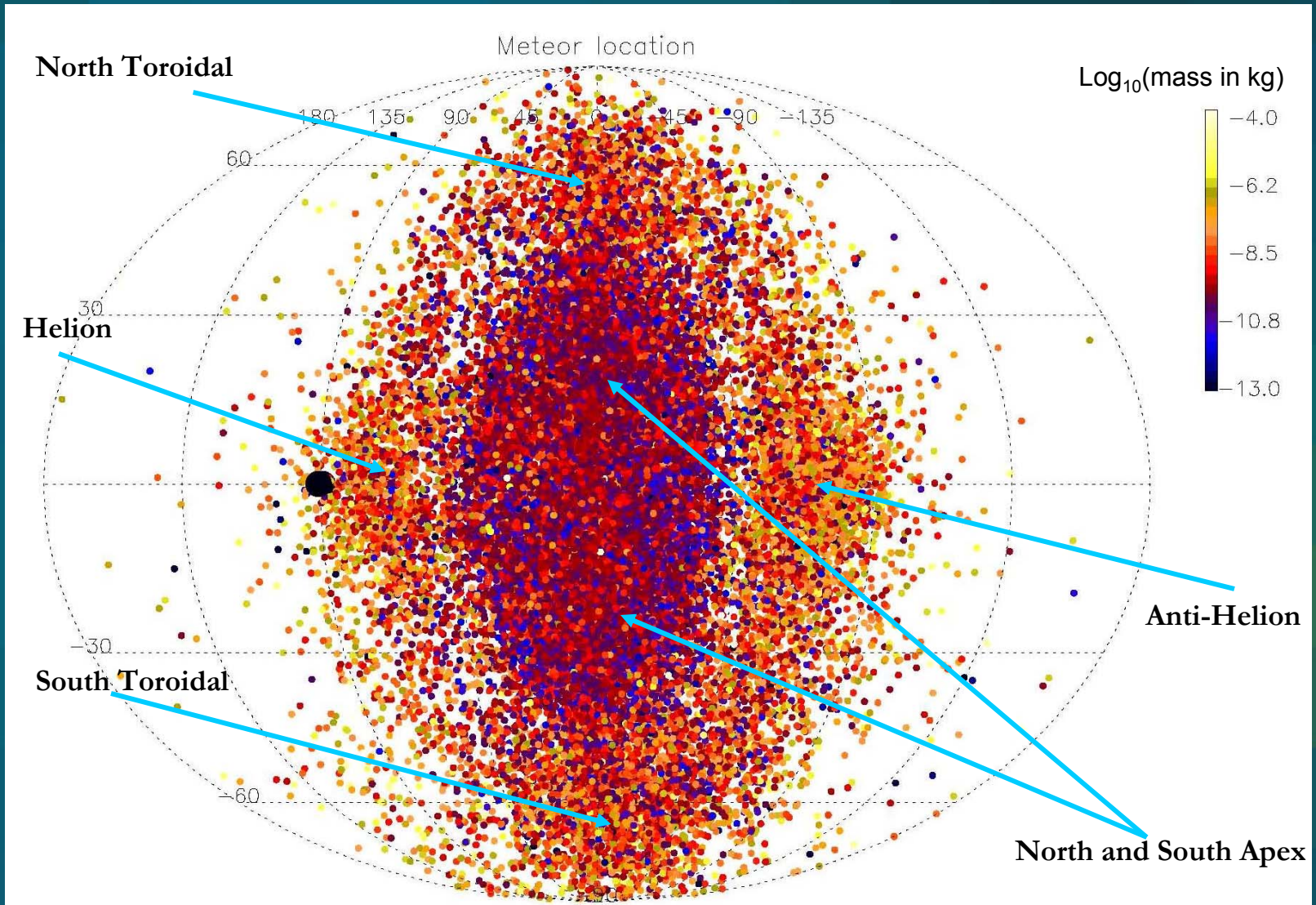
- 50 MHz radar at JRO
- Barker coded pulse
- Data collected over six days
- ~89,000 head echoes

Mass distribution

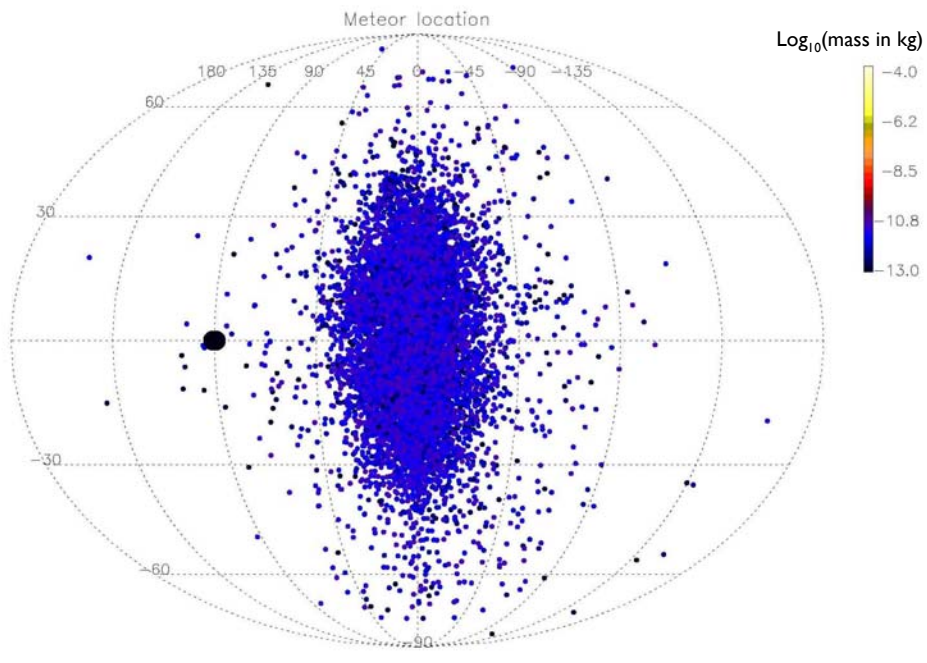


Spatial distributions

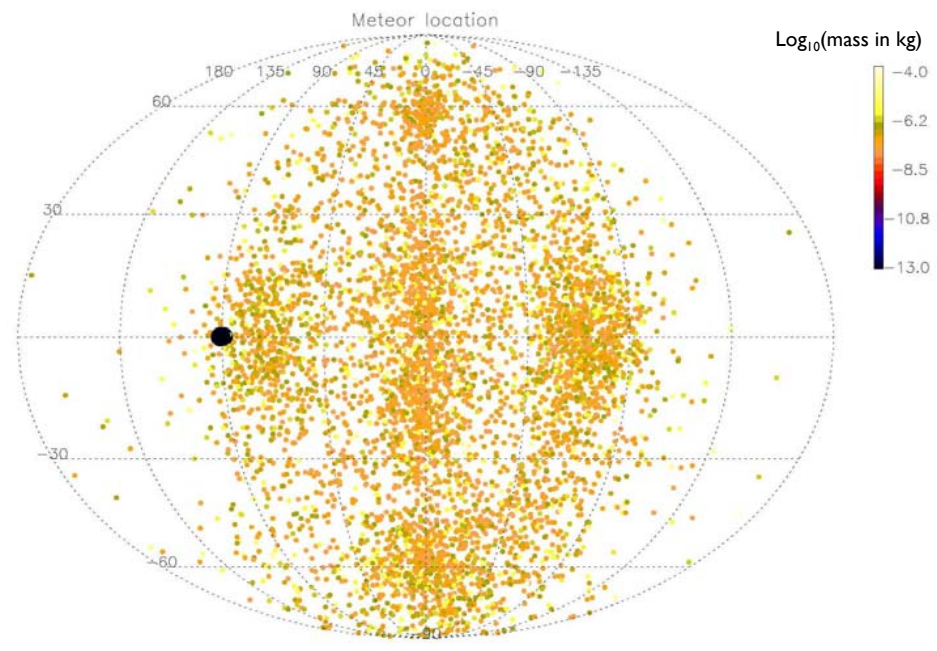
All Meteoroid Masses



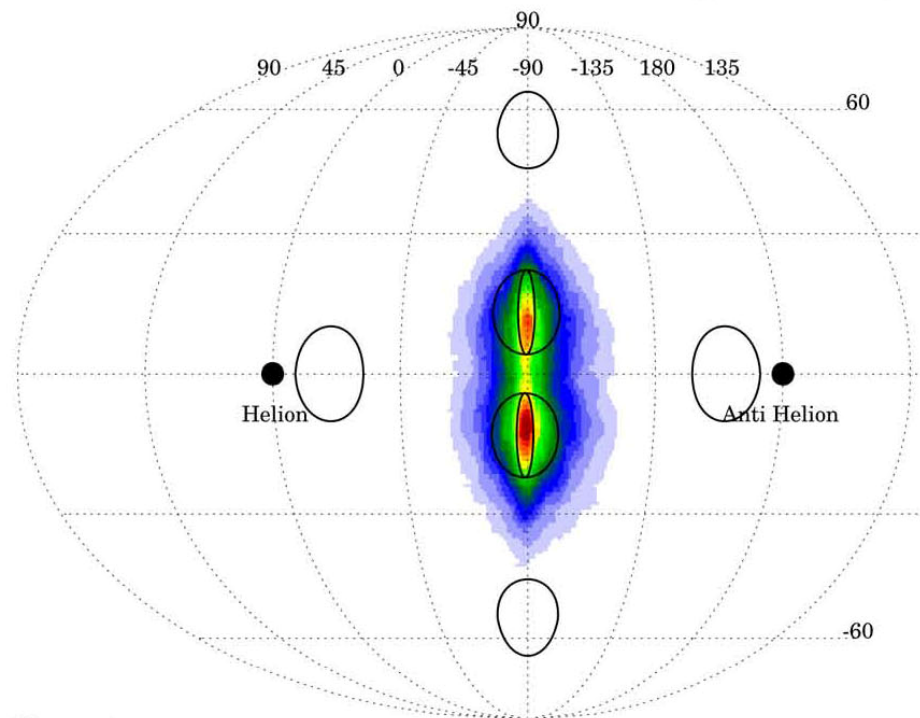
Meteoroids With Mass Less Than 10^{-12} kg



Meteoroids With Mass Greater Than 10^{-9} kg

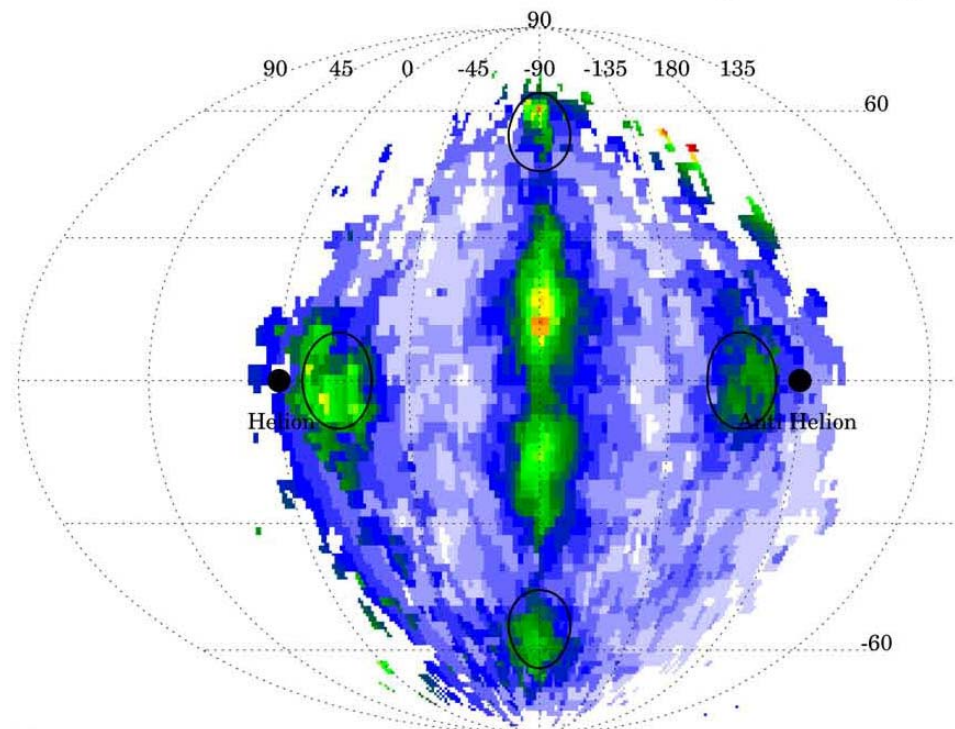


Distributions of all meteors before removing Earth velocity



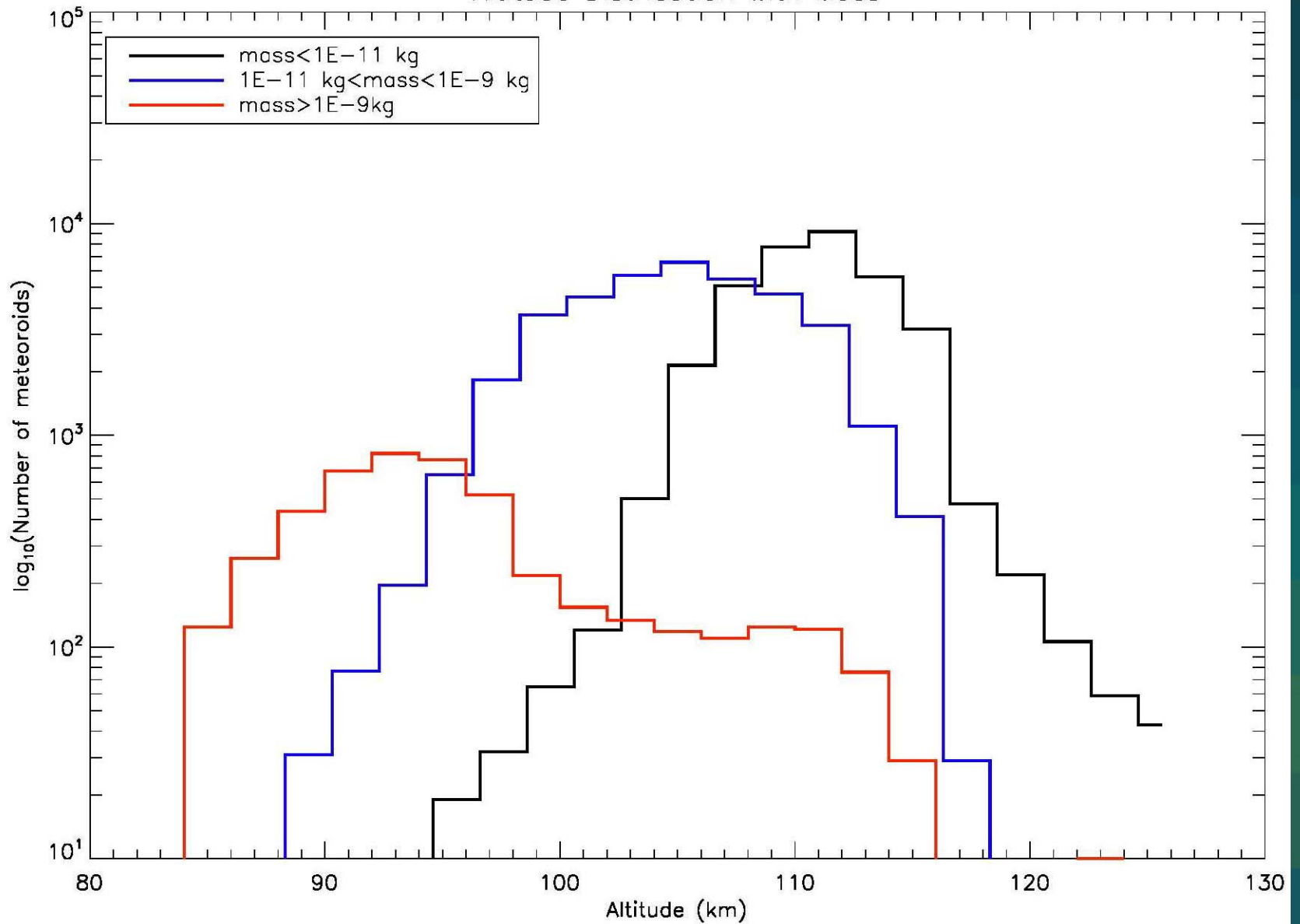
All campaigns

Distributions of all meteors before removing Earth velocity

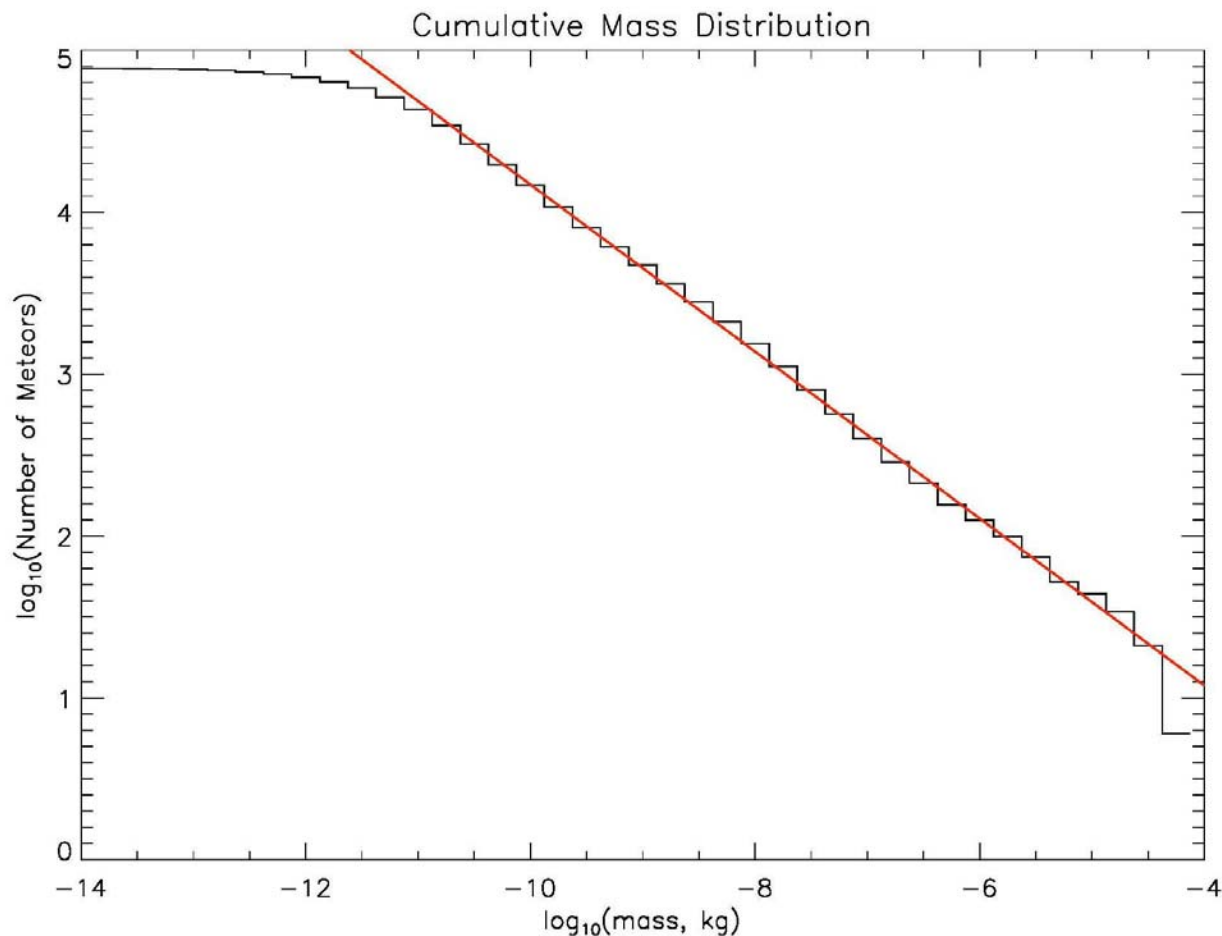


All campaigns

Altitude Distribution with Mass



Cumulative mass index



$$N(m > m_0) \propto m_0^{-\alpha}$$

- $\alpha = -0.515 \pm 0.004$
- Previous studies:
 - ALTAIR: $\alpha = 0.94$ (Close et al. 2007)
 - Photographic: $\alpha = 1.36$ (Whipple, 1967)
 - Specular: $\alpha = 1.0$ (Galligan and Baggaley, 2004)

Conclusions

- Meteoroids with smallest masses seen from Apex sources
- Larger meteoroids seen from all six sources
- Unusually low cumulative mass index found: $\alpha = 0.515$