

What else can we learn with coherent scatter radars?

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Acknowledgments

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Outline

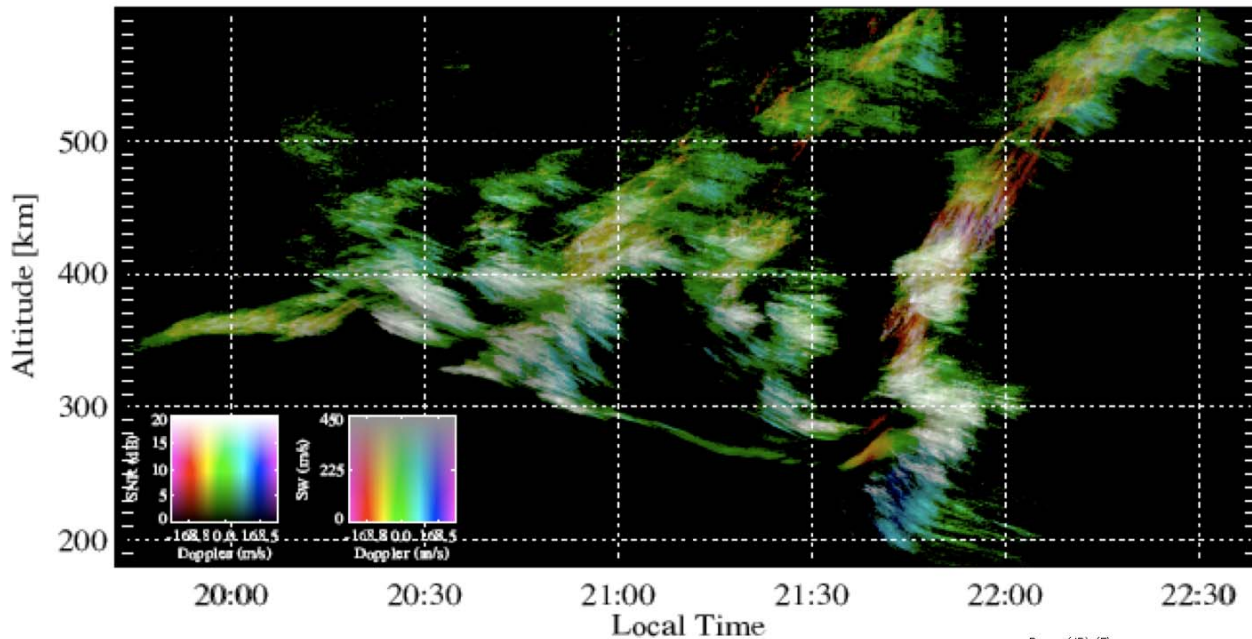
- ?
- Targets/Techniques/Plasma research vs. Diagnostics
(Declare victory?)
- 150-km echo challenge
- Mother Nature vs. Man Made
- Conclusions

Main “Equatorial” Questions

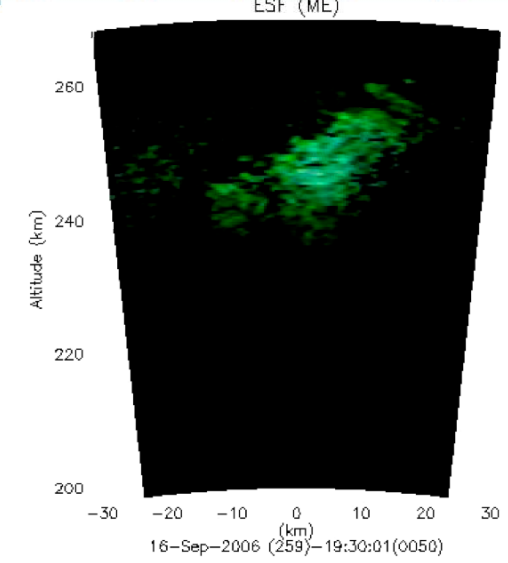
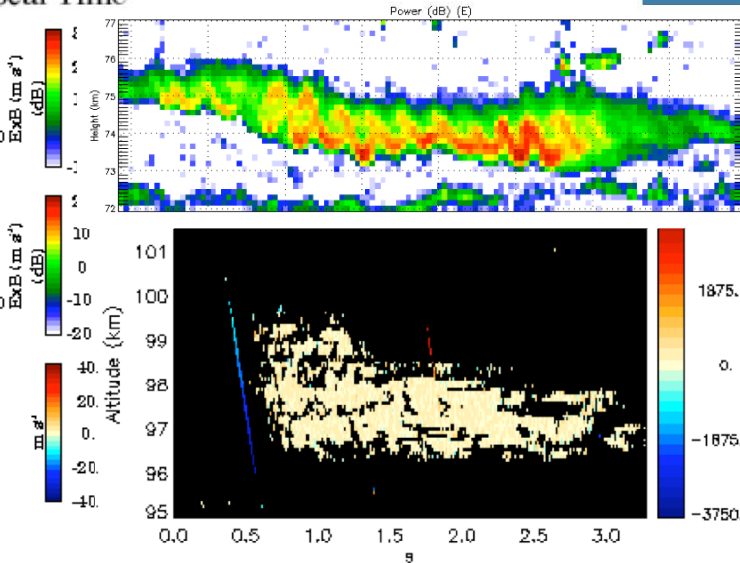
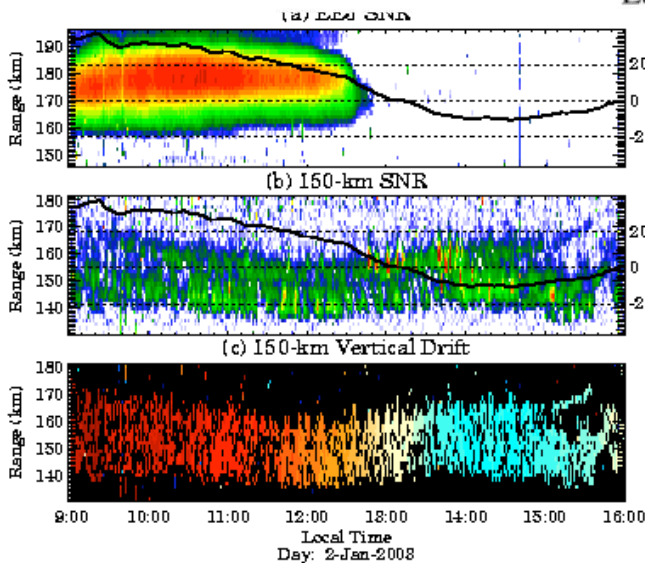
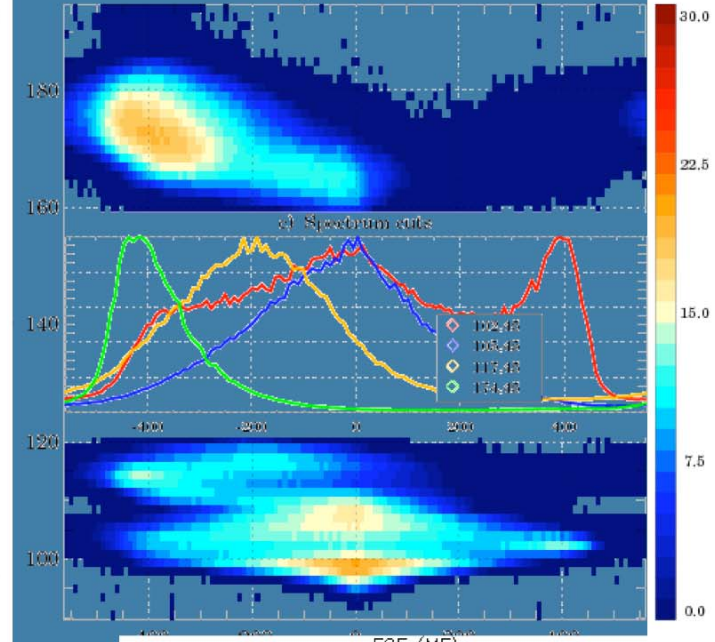
- *F* region: What are the **fundamental plasma processes**, including nonlinear processes, that govern the generation of plasma plumes? What are the **precursor phenomena** in the late afternoon *F* region that control whether or not an *F*-region plume will be generated after sunset?
- Daytime Valley echoes (or so-called 150-km echoes). What are the **physical mechanisms** causing them?
- *E* region: What are the **nonlinear plasma physics** processes that control the final state of the electrojet instabilities? To what extent do these instabilities **affect the conductivity** of the *E* region.
 - What are the basic background parameters in the equatorial *E* region? What is the morphology of the density profiles in this difficult to probe region? How does this morphology affect the *E*-region dynamo?
- *D* region: What effects do meteor ablation and mesospheric mixing have on the composition in this region?
- *E* and *F* (valley) region coupling. Does the *F* region respond to an **Es layer instability**? Are **150-km echoes** related to *Es* layers?

Atmospheric/Ionospheric Irregularities

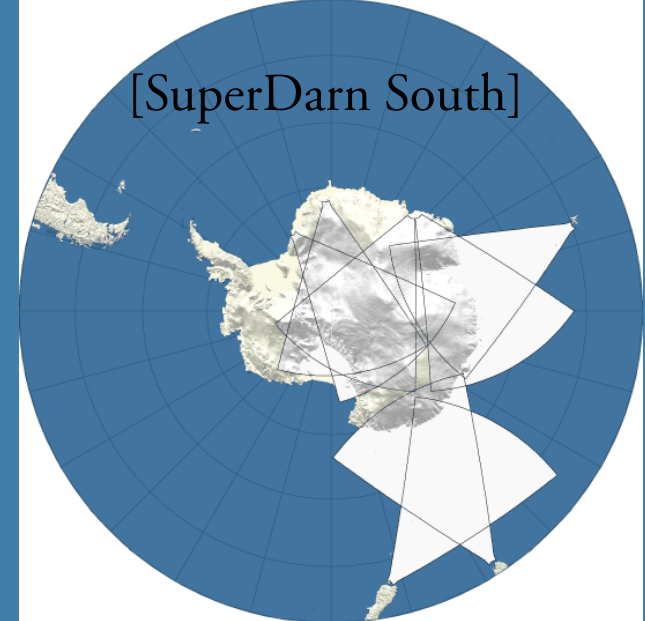
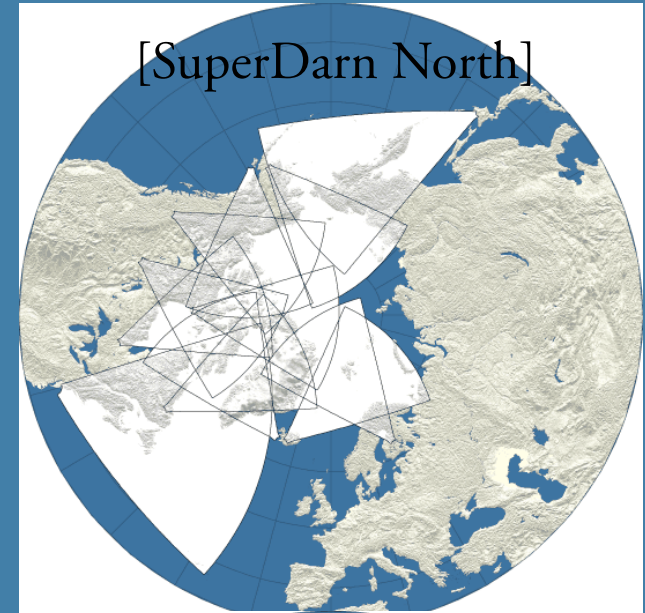
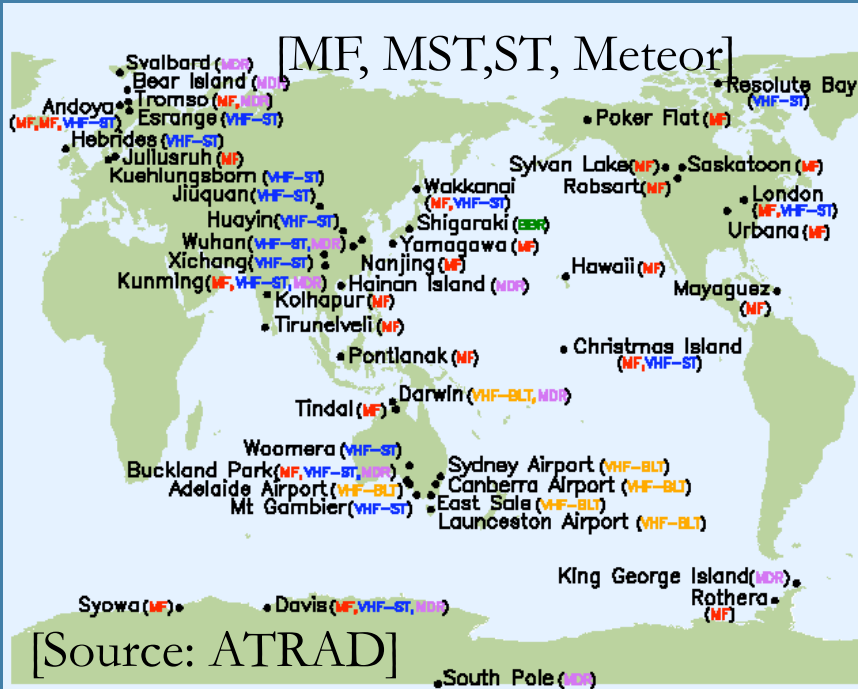
RTDI over JRO



a) 11:02:09



Coherent Radar Networks



Coherent Radars Summary

Radar Frequency

MF

HF

VHF

UHF

Configurations/ Techniques

Monostatic/Multi-Static

Multi-beam

CW/Pulsed

Multi-station

Interferometry

Imaging

Multi Frequency

Passive

Coherent Targets

PEME/PMSE

Specular Meteors

Other meteors

E region

150-km/Valley

F region

Main Derived Parameters

Irregularity Power,
Drifts, and spectra shape

Neutral winds

Electric fields

Electron Density

Neutral Temperatures

GWs/Tides/PWs

Coherent Radars: SuperDarn

Radar Frequency

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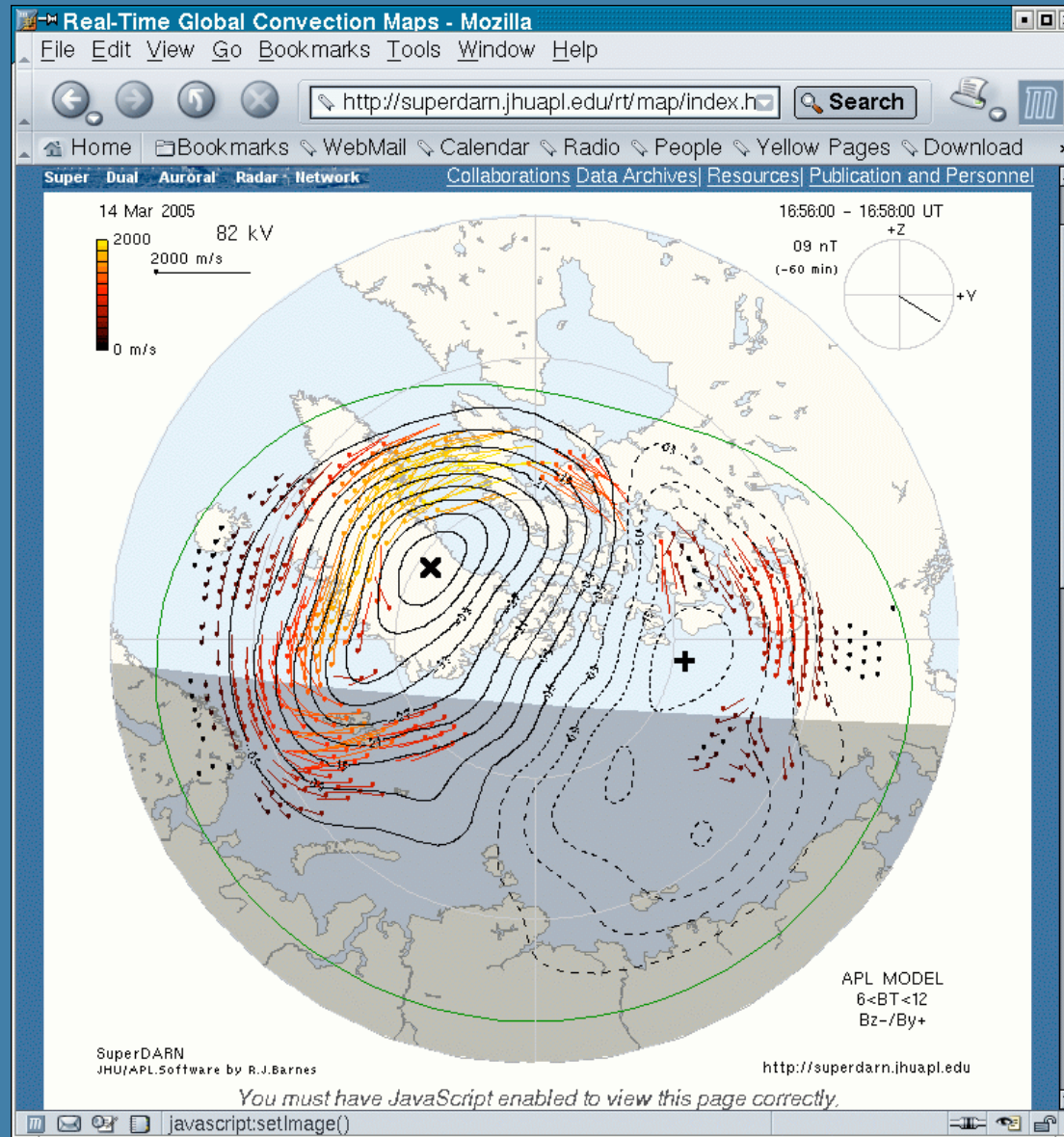
Electric fields

Electron Density

Neutral Temperatures

GWs/Tides/PWs

SuperDarn Convection Maps



Coherent Radars: ESF Imaging

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VHF

UHF

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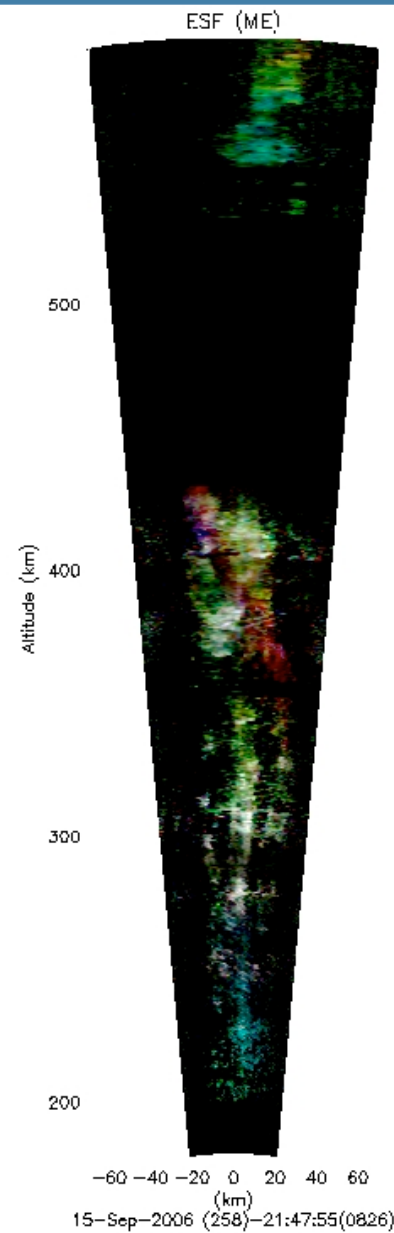
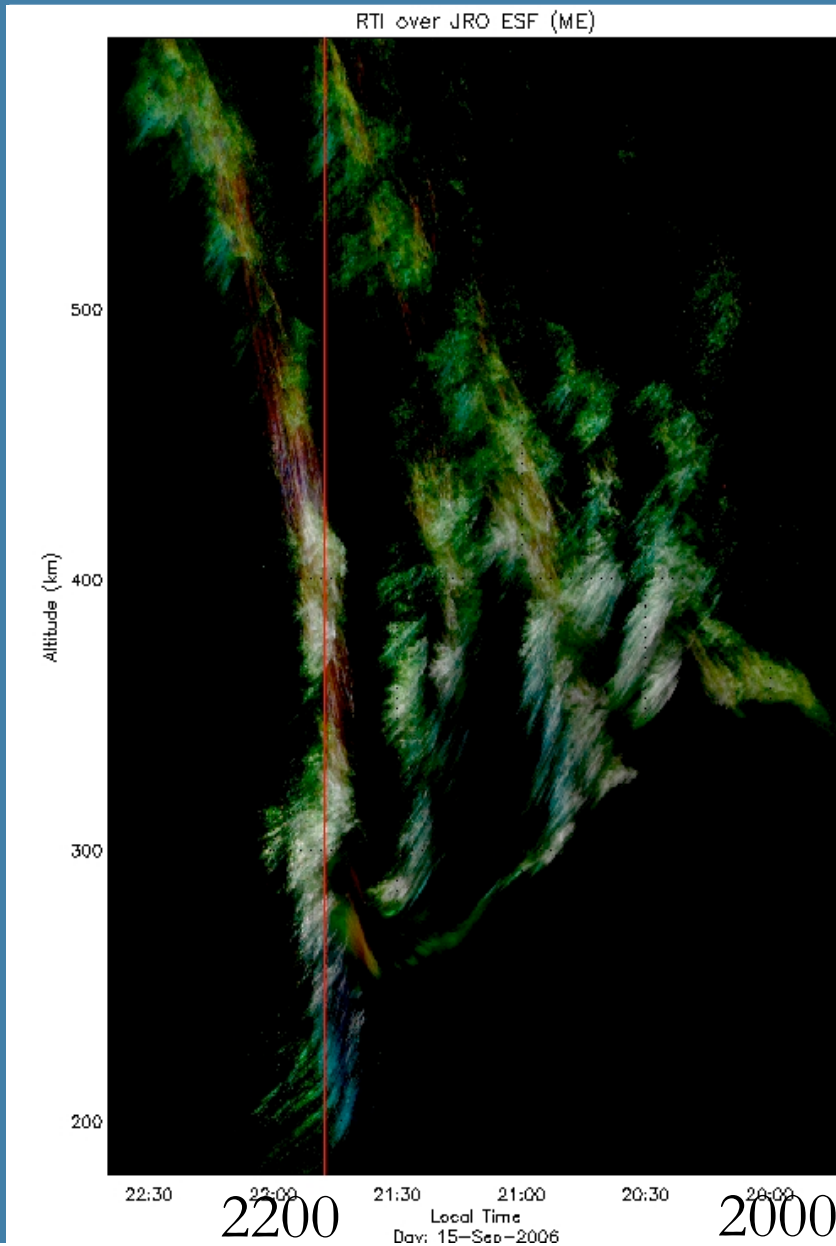
ESF RTDI + Imaging

500

400

300

200



[Chau et al., 2008]

Coherent Radars: 150-km Echoes

Radar Frequency

MF

HF

VHF

UHF

Configurations/ Techniques

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Multi-beam

CW/Pulsed

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Interferometry

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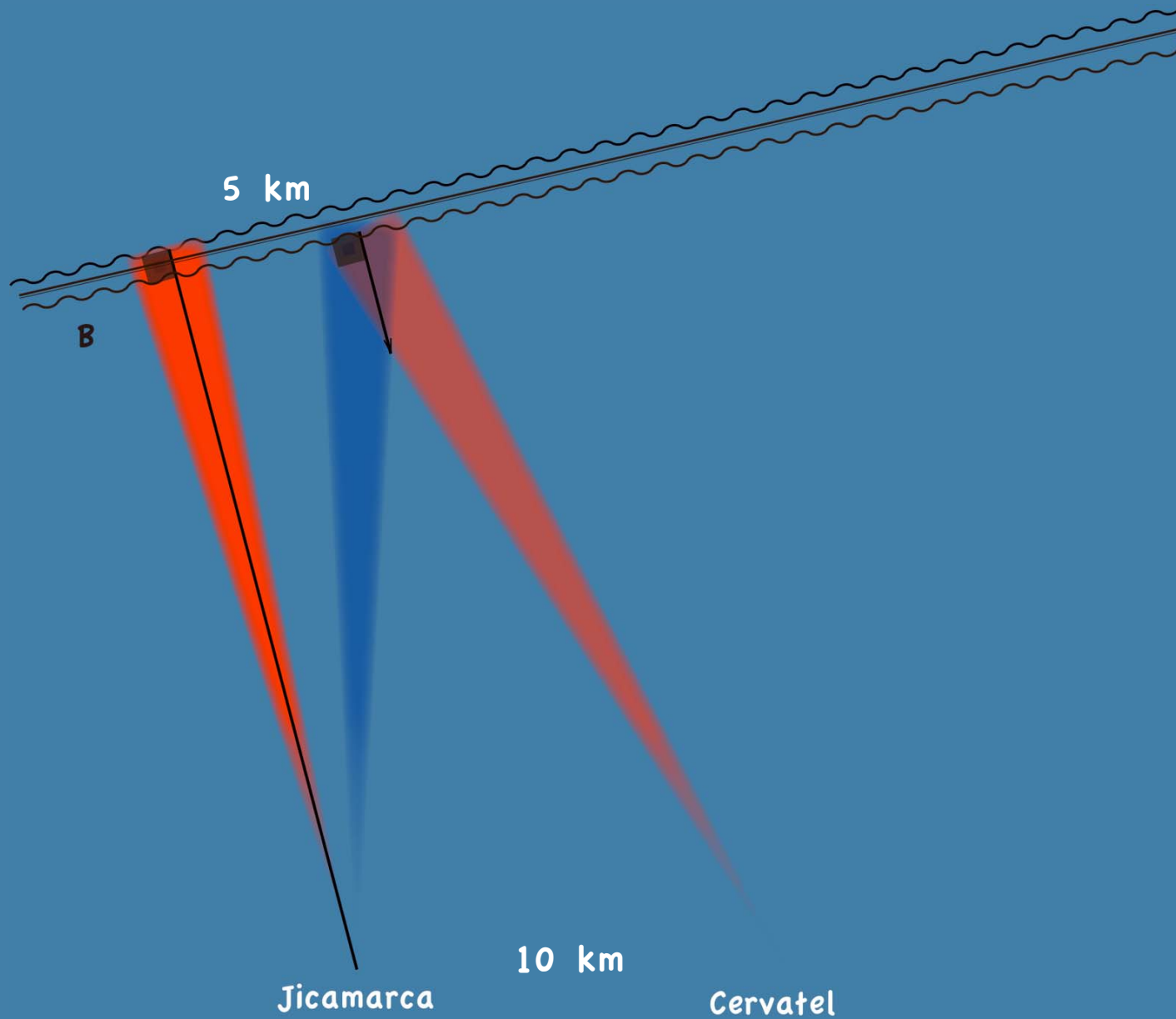
Electric fields

Electron Density

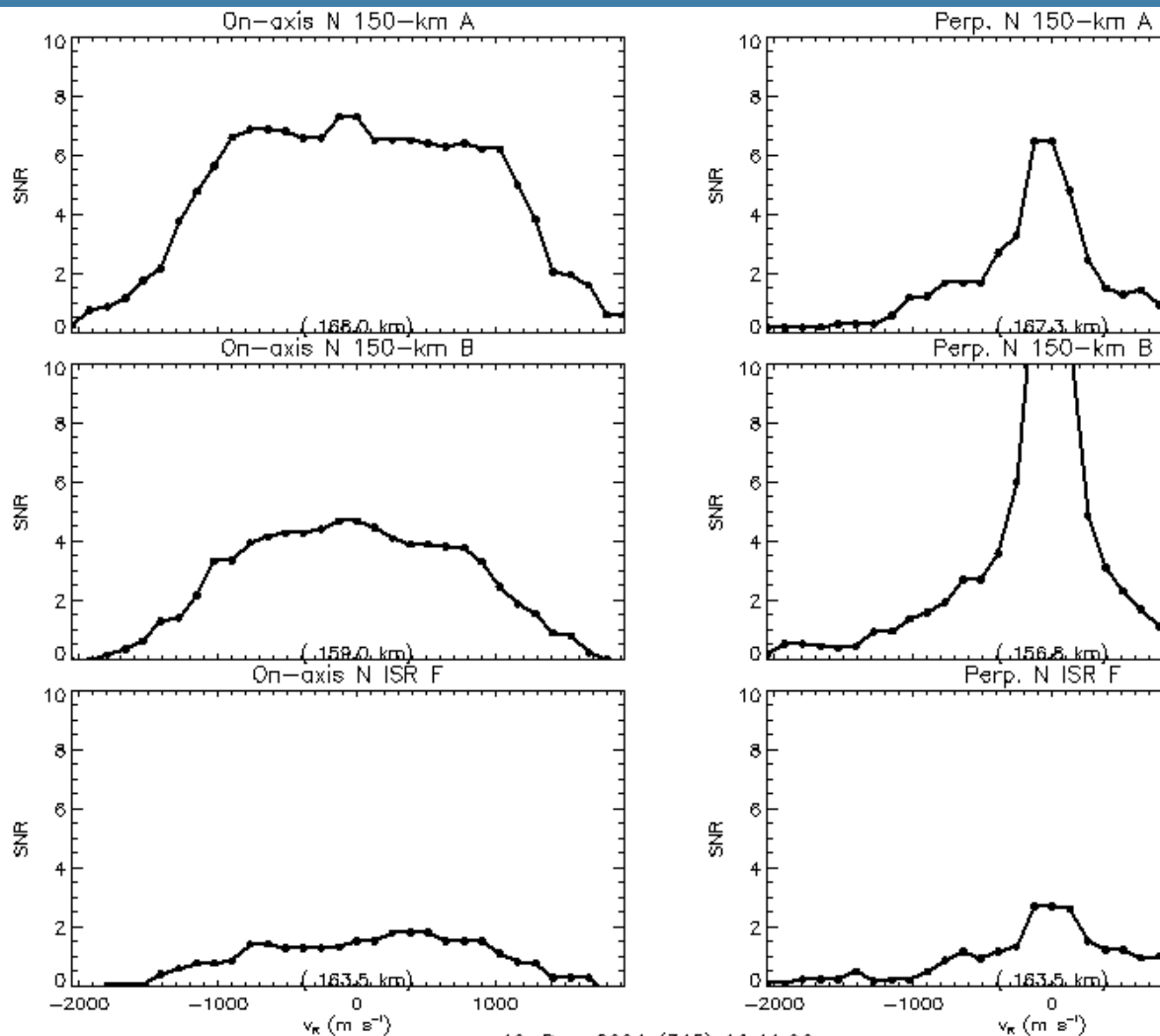
Neutral Temperatures

GWs/Tides/PWs

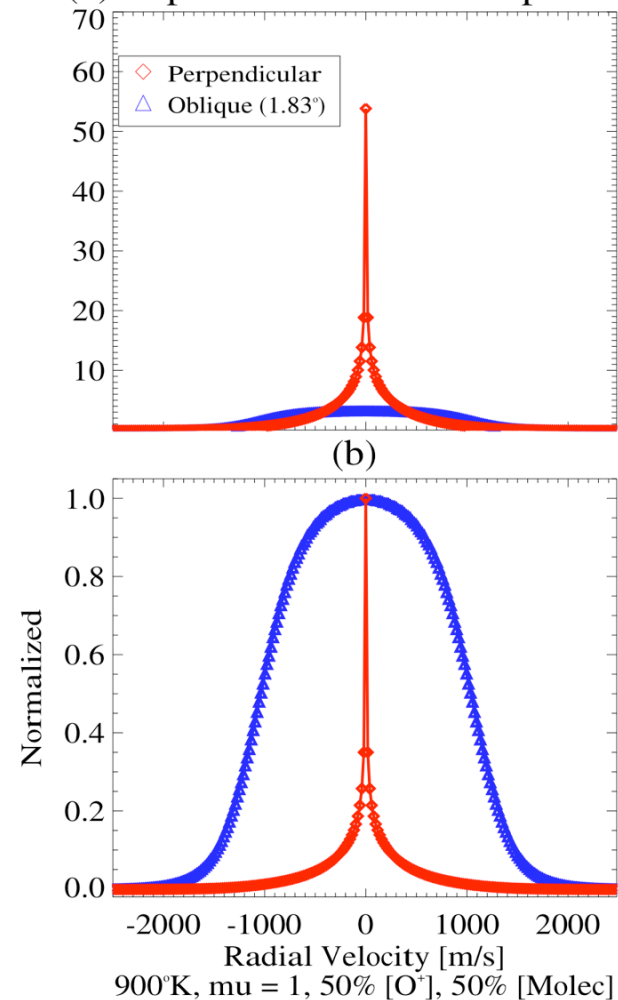
150-km Experiments: Oblique vs. Perpendicular



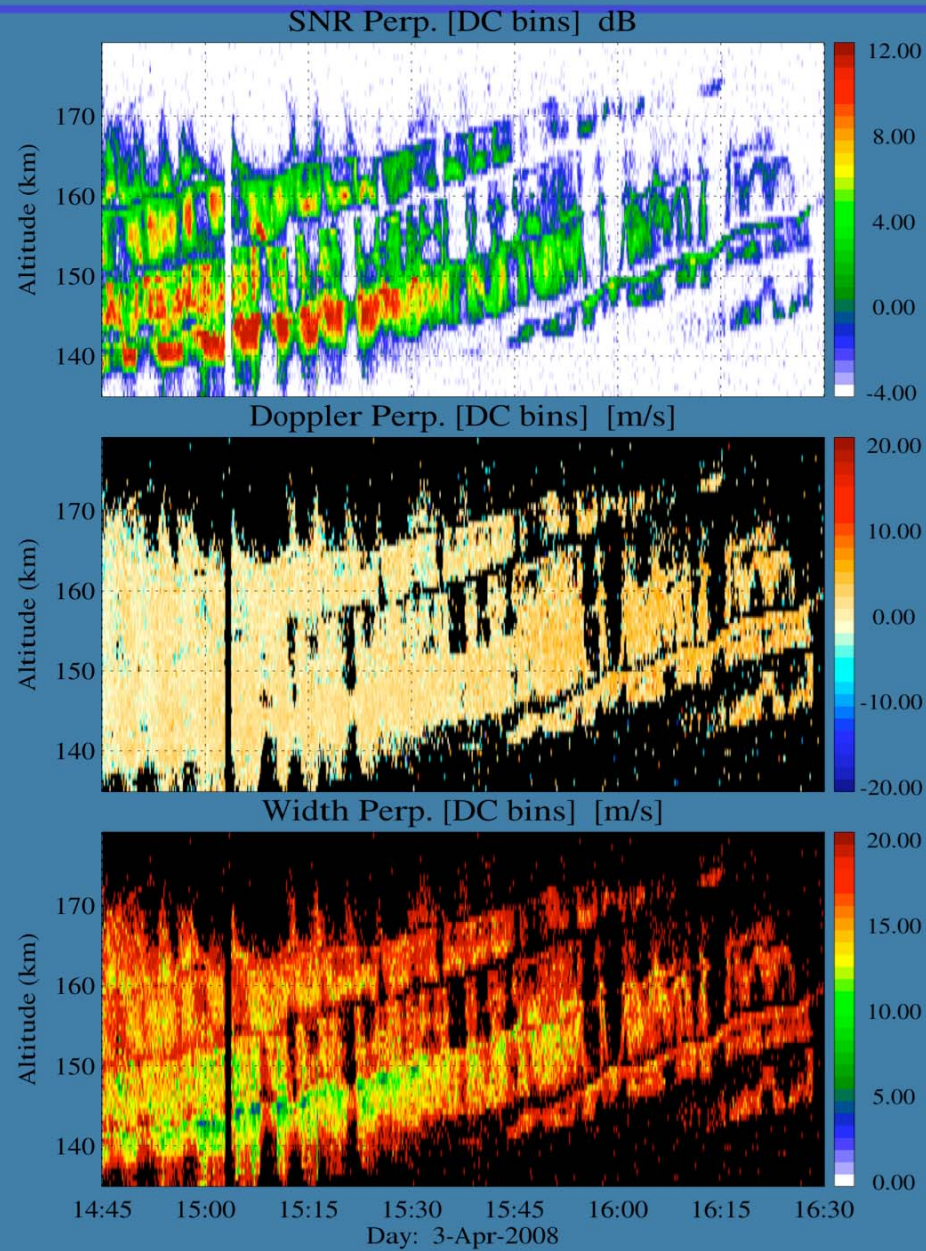
150-km Spectra: Oblique vs. Perpendicular



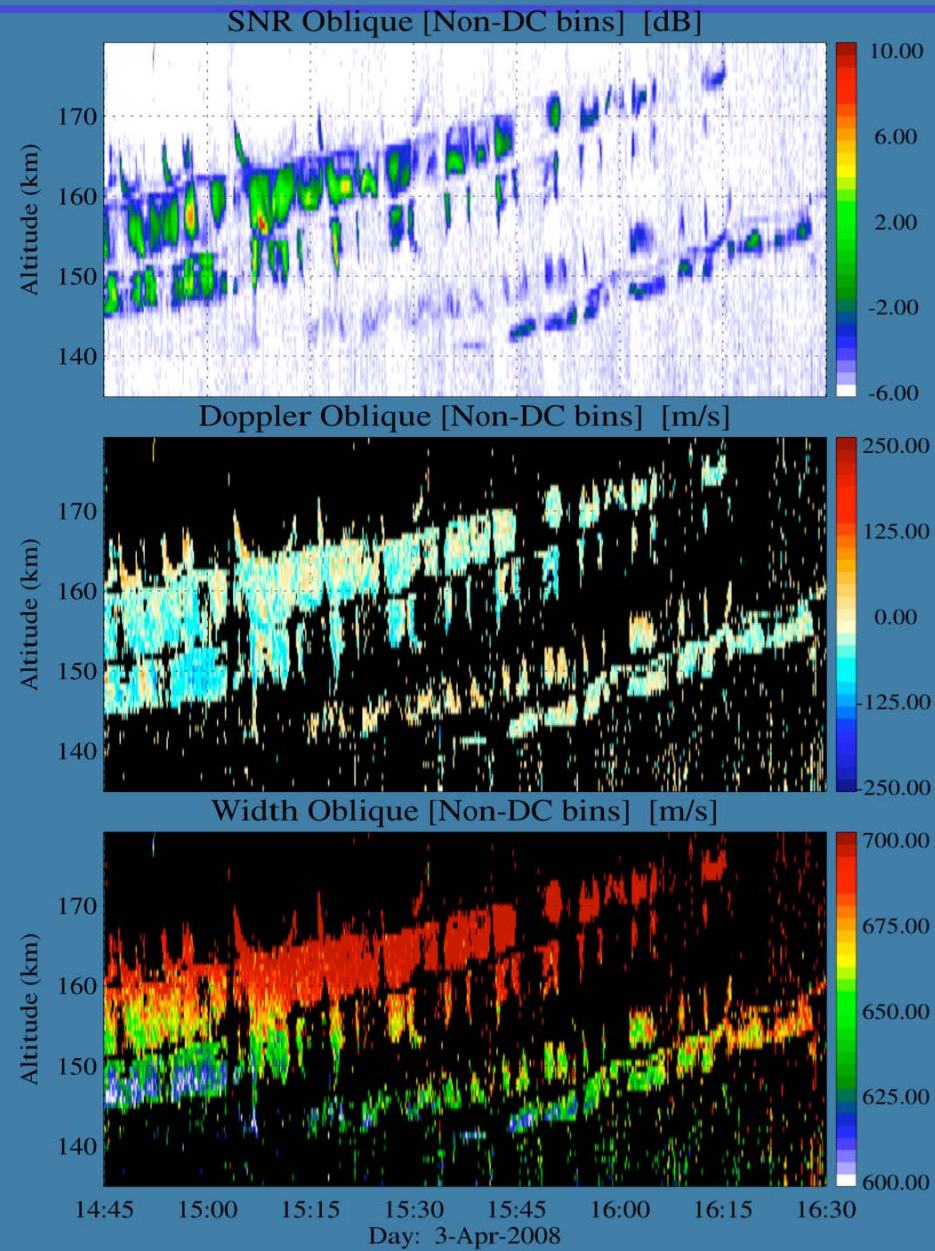
(a) Expected 150-km ISR Spectrum



150-km Perpendicular Parameters

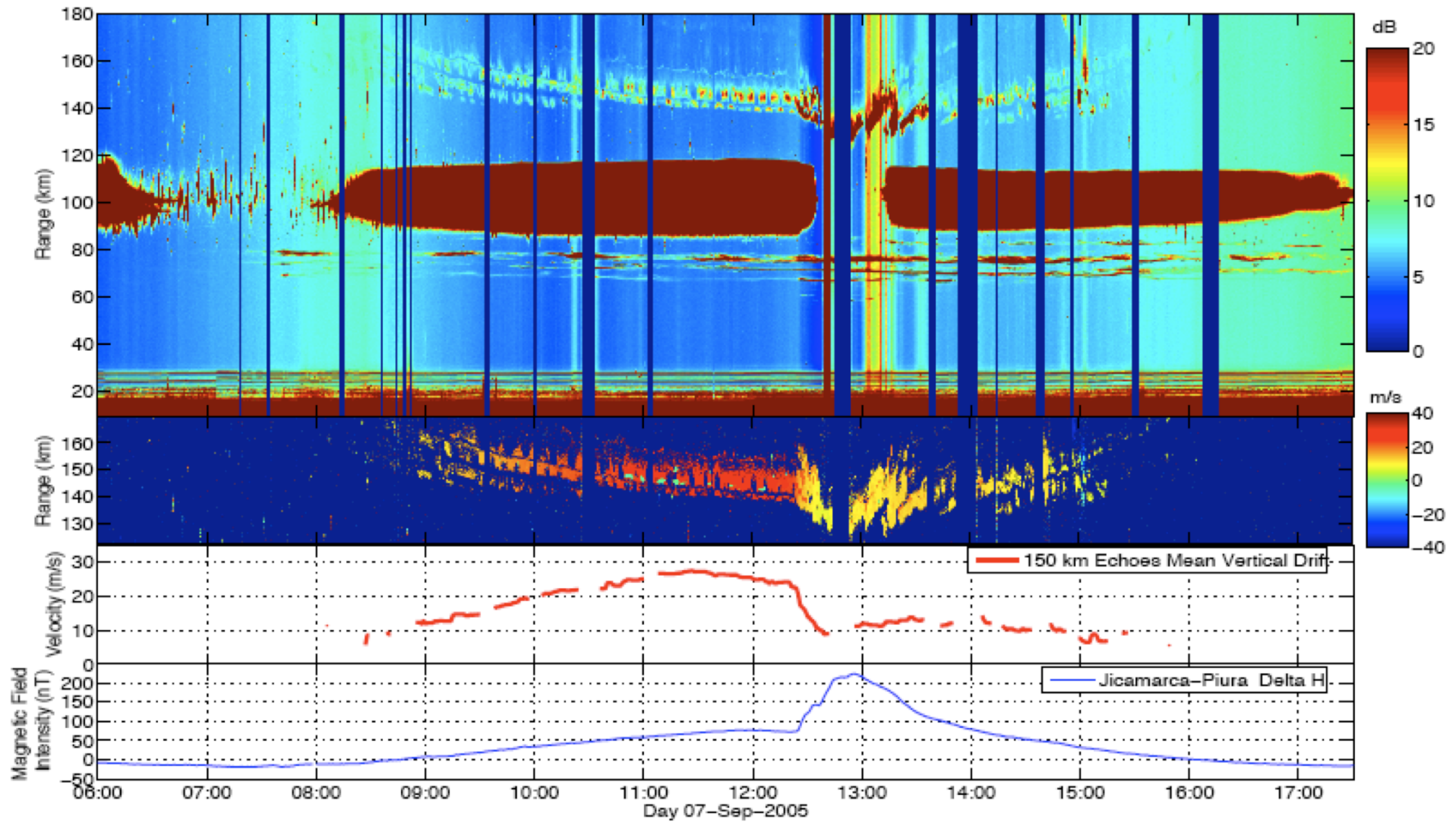


150-km Oblique Parameters



Equatorial Irregularities modified by Solar Flares

East Beam RTI & 150 km vertical Drifts



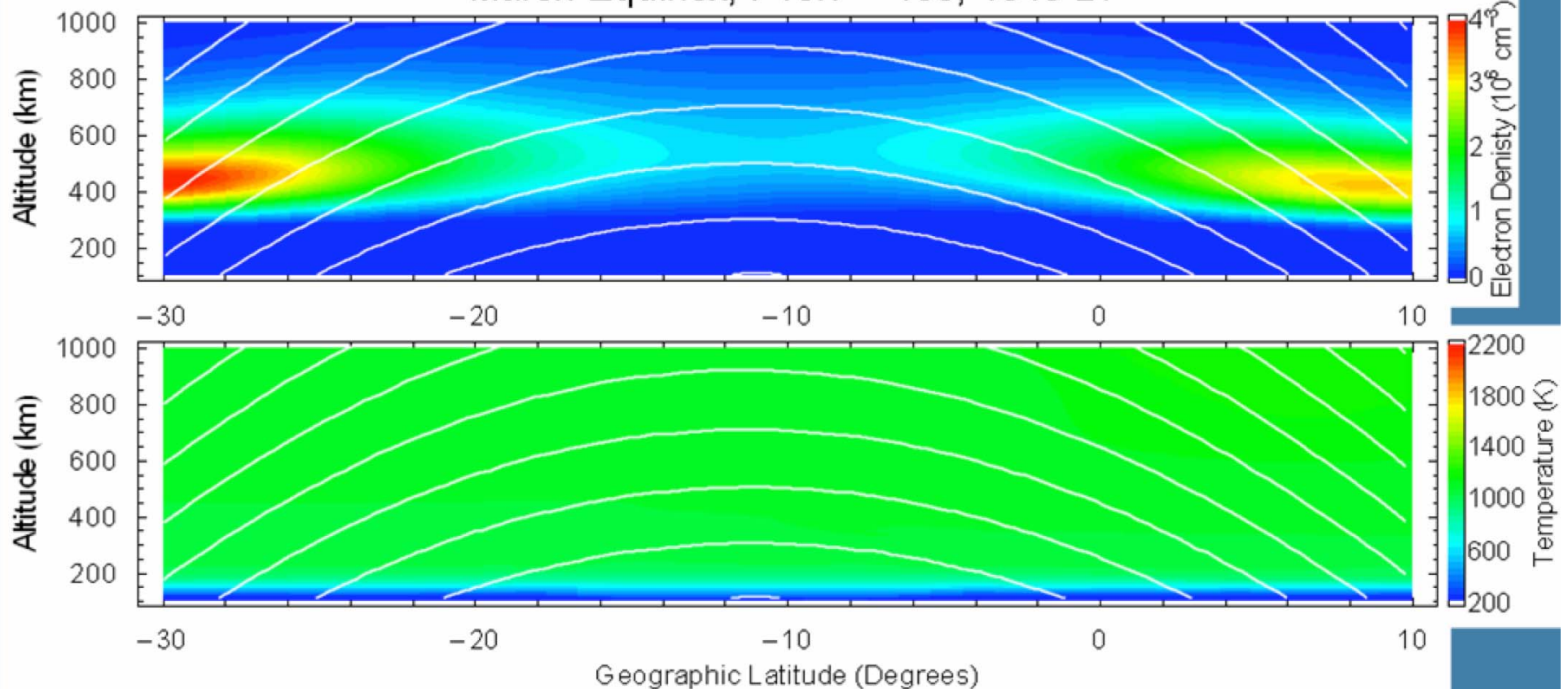
Solar flare 07-Sep-2005

[Courtesy of P. Reyes]

SAMI2 Model

283° Longitude Equatorial Ionosphere

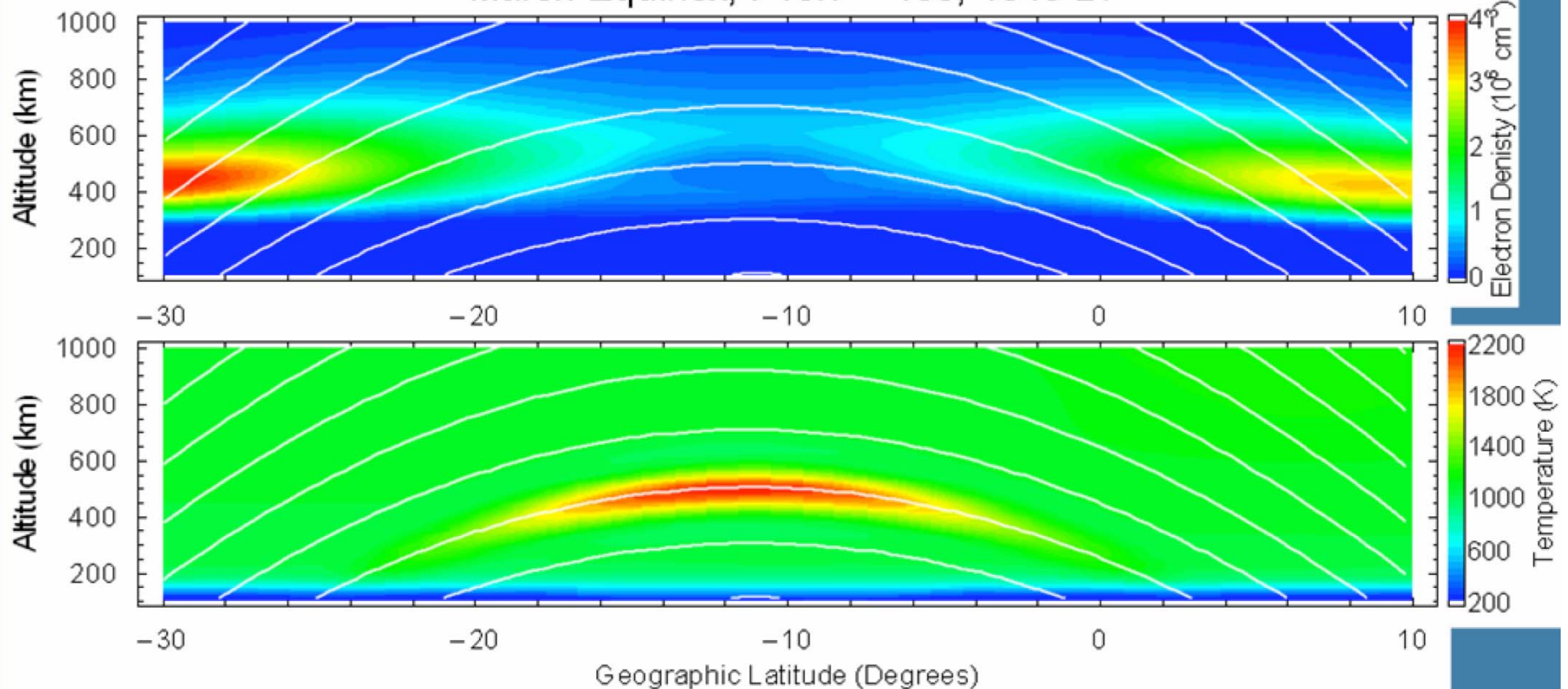
March Equinox, F10.7 = 150, 1845 LT



[Courtesy of P. Bernhardt]

Heated Field Line at 7.9 MHz ($7.8 \times 10^5 \text{ cm}^{-3}$) 283° Longitude Equatorial Ionosphere

March Equinox, F10.7 = 150, 1845 LT



[Courtesy of P. Bernhardt]

Concluding Remarks

- Irregularity drifts vs. electric fields (local and background)
- Radar Imaging.
- Common volume multi-frequency.
- E and F (valley) region coupling.
- 150-km campaigns (multi-instrument?)
- Coherent scatter diagnostics of artificial ionospheric irregularities at equatorial and low latitudes.